

# 利用GeV能区重离子碰撞研究高密核物质

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## 二、CEE (HIRFL-CSR) 和 CEE+(HIAF) 的新机遇

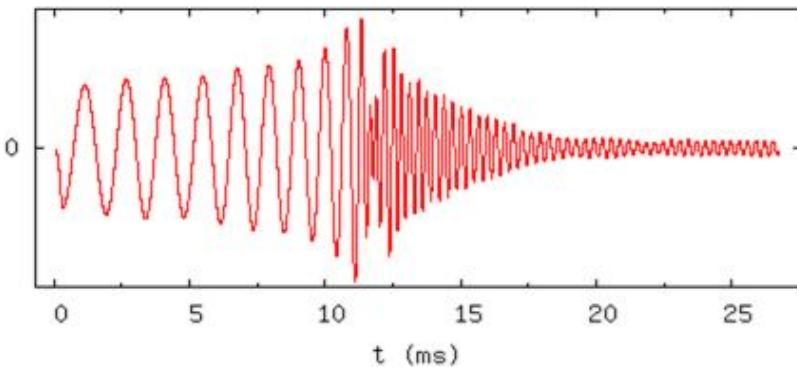
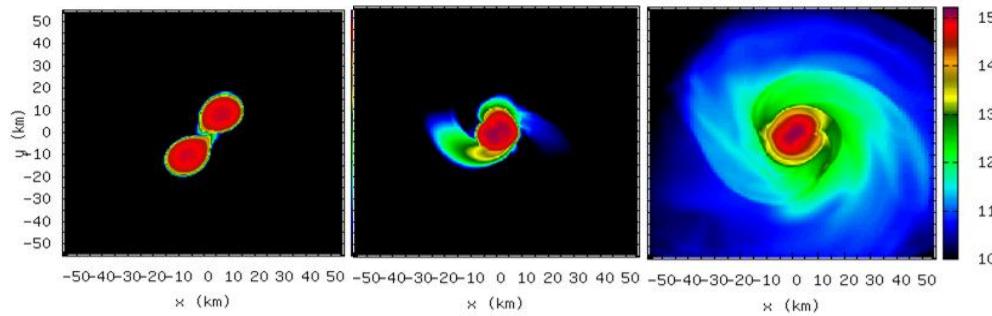
1. CEE: EOS物理目标可行性
2. HIAF: 更高密度区的核物质状态方程

## 三、小结

# 双中子星并合事件及其核物理相关性

## GW170817 及其与核物质状态方程的关联

$t = 10 \text{ ms}$      $t = 11.2 \text{ ms}$      $t = 13.9 \text{ ms}$



$$E_{\text{sym}}(\rho) \longleftrightarrow \text{R-M 关系} \longleftrightarrow \text{GW}$$

TOV Eq.

$\Lambda \propto (R/M)^5$

## TOV 方程

$$\frac{d\mathcal{M}(r)}{dr} = 4\pi r^2 \varepsilon(r), \quad -\frac{d\mathcal{P}(r)}{dr} = \frac{G e \mathcal{M}}{r^2} \left(1 - \frac{2G\mathcal{M}}{r}\right)^{-1} \left(1 + \frac{\mathcal{P}}{\varepsilon}\right) \left(1 + \frac{4\pi r^3 \mathcal{P}}{\mathcal{M}}\right)$$

# 核物质状态方程

The nucleon specific energy of nuclear matter formed in HIC or in neutron star can be written as

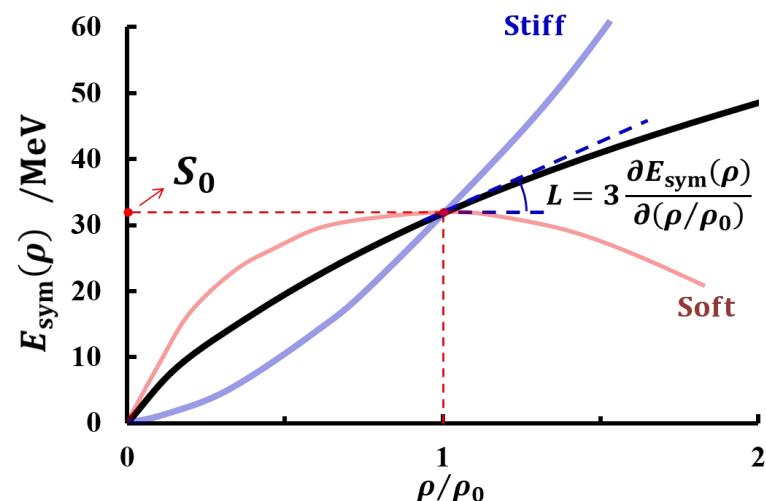
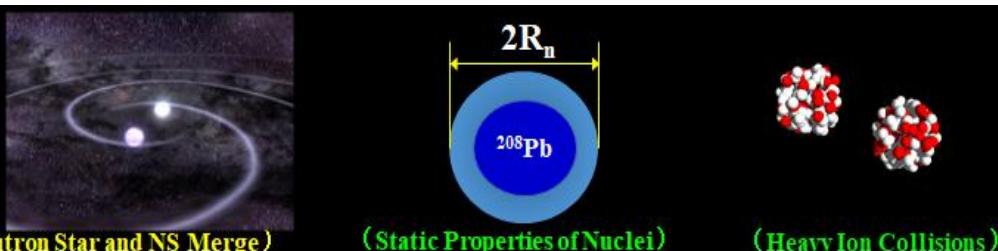
$$E(\rho, \delta) = E_0(\rho) + \delta^2 E_{\text{sym}}(\rho), \quad \delta = \frac{\rho_n - \rho_p}{\rho_n + \rho_p}$$

$$E_0(\rho) = E_0(\rho_0) + \frac{K_0}{2} \left( \frac{\rho - \rho_0}{3\rho_0} \right)^2 + \frac{J_0}{6} \left( \frac{\rho - \rho_0}{3\rho_0} \right)^3 + \dots$$

$$E_{\text{sym}}(\rho) = E_{\text{sym}}(\rho_0) + L \left( \frac{\rho - \rho_0}{3\rho_0} \right) + \frac{K_{\text{sym}}}{2} \left( \frac{\rho - \rho_0}{3\rho_0} \right)^2 + \frac{J_{\text{sym}}}{6} \left( \frac{\rho - \rho_0}{3\rho_0} \right)^3 + \dots$$

For npe matter, the pressure :

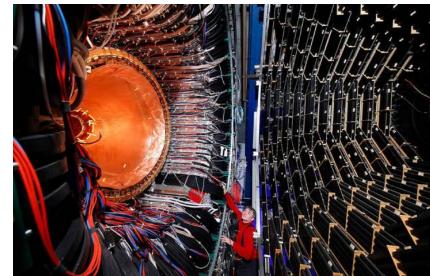
$$P(\rho, \delta) = \rho^2 \frac{\partial E(\rho, \delta)}{\partial \rho} + \frac{1}{4} \rho e \mu_e = \rho^2 \left[ \frac{\partial E_0}{\partial \rho} + \delta^2 \frac{\partial E_{\text{sym}}}{\partial \rho} \right] + \frac{1}{2} \delta(1 - \delta) \rho E_{\text{sym}}(\rho)$$



# 世界几个主要的GeV能区重离子实验

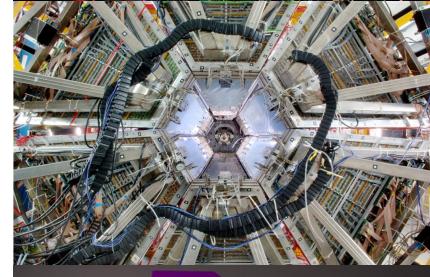
## 1. Plastic Ball at Bevalac

PRC 42, 640 (1990) ...



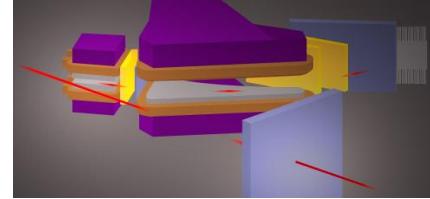
## 2. EOS at Bevalac

PRL 78, 2535 (1997); RPC 76, 3911 (1996); PRL 75, 2662 (1995) ...



## 3. FOPI at SIS

NPA 876, 1 (2012); NPA 848, 366 (2010); NPA 781, 459 (2007) ...



## 4. KAOS at SIS

PRL 96, 072301 (2006); PRL 95, 012301 (2005); PRL 86, 1974 (2001) ...

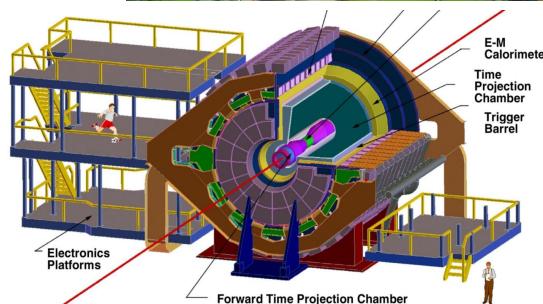


## 5. HADES at SIS

PRL 125, 262301 (2020); PRL 123, 022002 (2019); PRC 102, 024914 (2020) ...

## 6. S<sub>π</sub>RIT at RIBF

PRL 126, 162701 (2021); PLB 822, 136681 (2021); PLB 813, 136016 (2021) ...

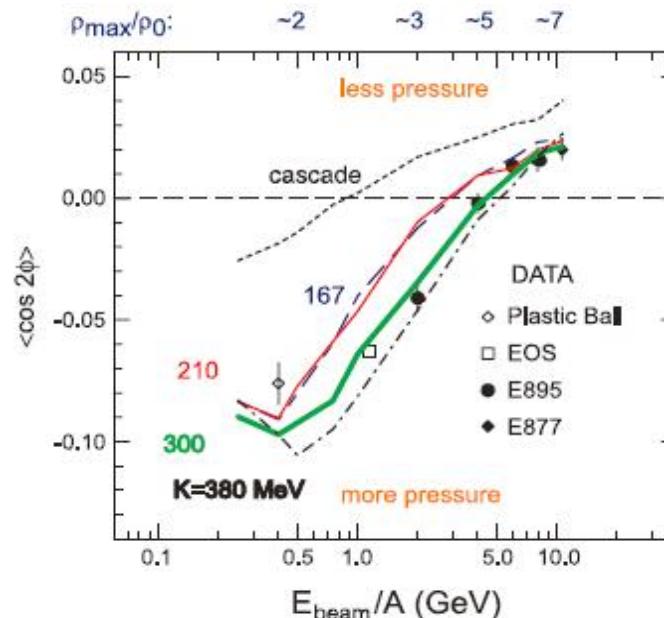
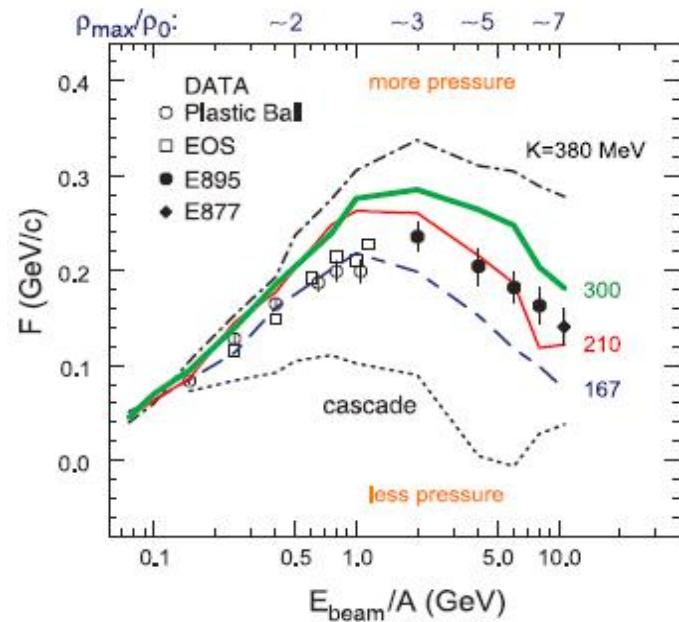


## 7. STAR-BES at RHIC

PRL 128, 202303 (2022); PLB 827, 137503 (2022) ...

and a lot more literatures...

# 通过集体流( $v_1, v_2$ )提取核物质状态方程



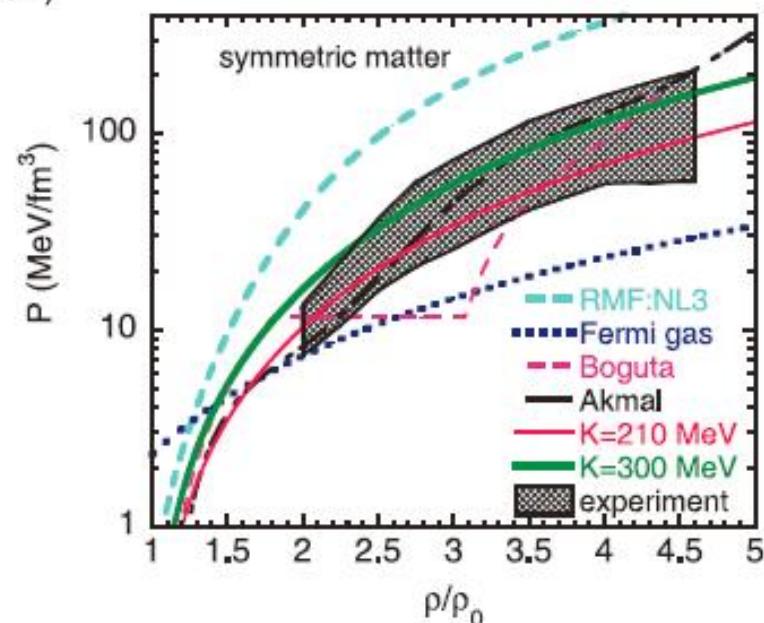
当两个原子核碰撞，会造成核物质的压缩，形成压力梯度，出现挤压效应，从而产生集体运动，形成所谓的集体流！

$$\frac{dN}{u_t du_t dy d\phi} = v_0 [1 + 2v_1 \cos(\phi) + 2v_2 \cos(2\phi)]$$

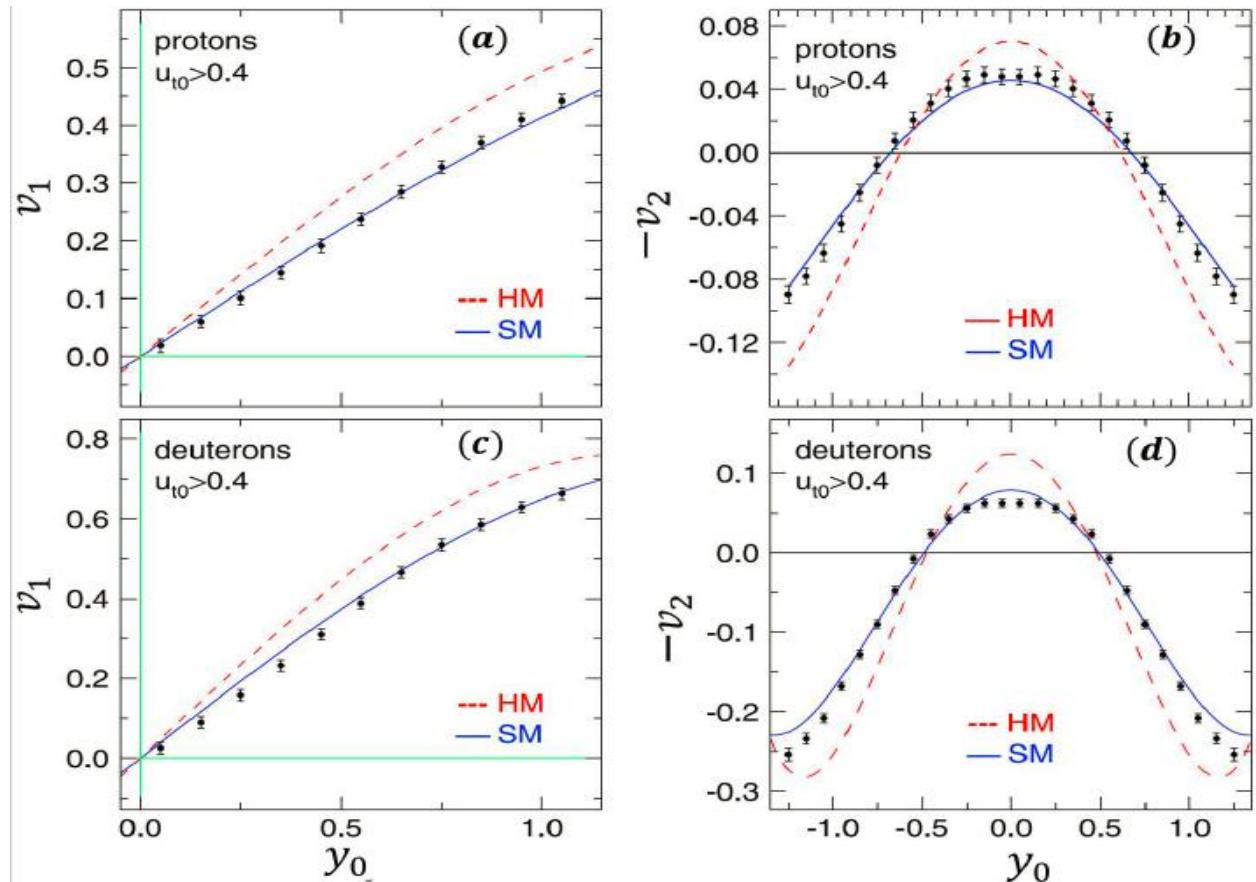
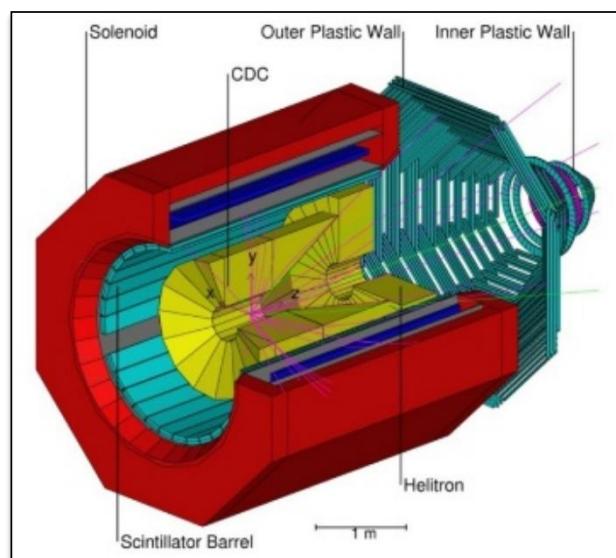
$$v_1 = \left\langle \frac{p_x}{p_t} \right\rangle = \langle \cos(\phi) \rangle,$$

$$v_2 = \left\langle \left( \frac{p_x}{p_t} \right)^2 - \left( \frac{p_y}{p_t} \right)^2 \right\rangle = \langle \cos(2\phi) \rangle$$

P. Danielewicz, R. Lacey, and W. G. Lynch.  
Determination of the equation of state of dense matter.  
Science, 298, 1592 (2002)



# 集体流( $v_1, v_2$ )提取核物质状态方程

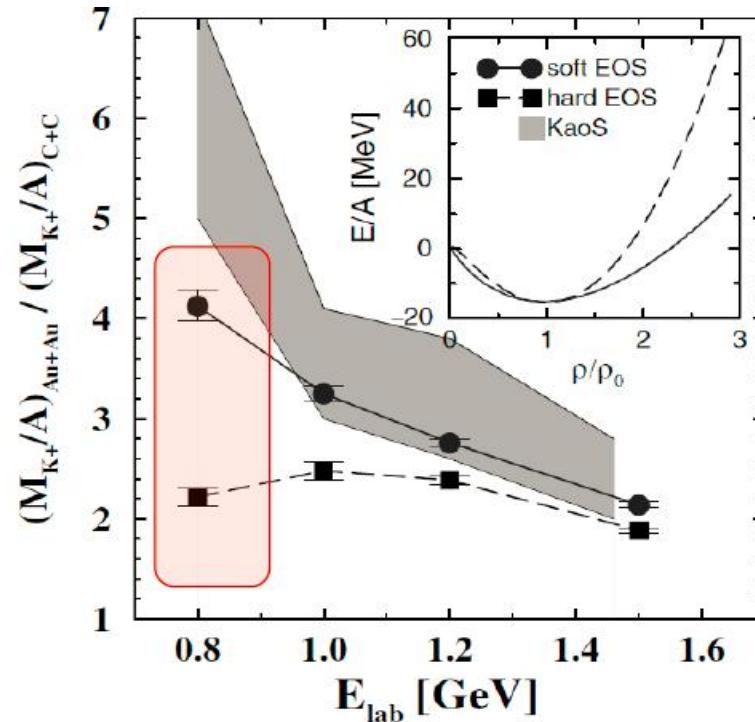
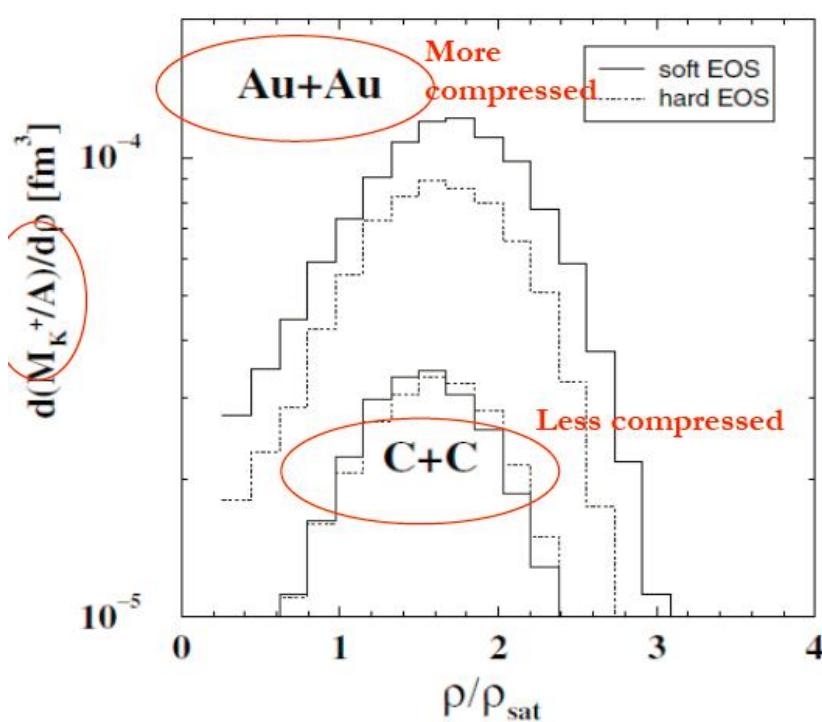


$$v_i = v_i(u_t, y)$$

Systematics of azimuthal asymmetries in heavy ion collisions in the 1 AGeV regime  
W. Reisdorf et al for FOPI collaboration, NPA 876, 1 (2012)

# K<sup>+</sup>产额提取核物质状态方程

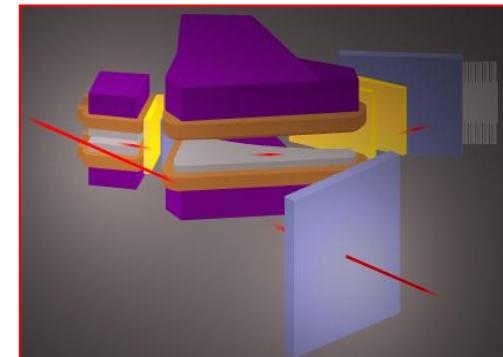
除了集体流之外，介子产额也和其产生时的介质性质密切相关！



J. Aichelin and Che Ming Ko, **PRL** 55,2661 (1985)

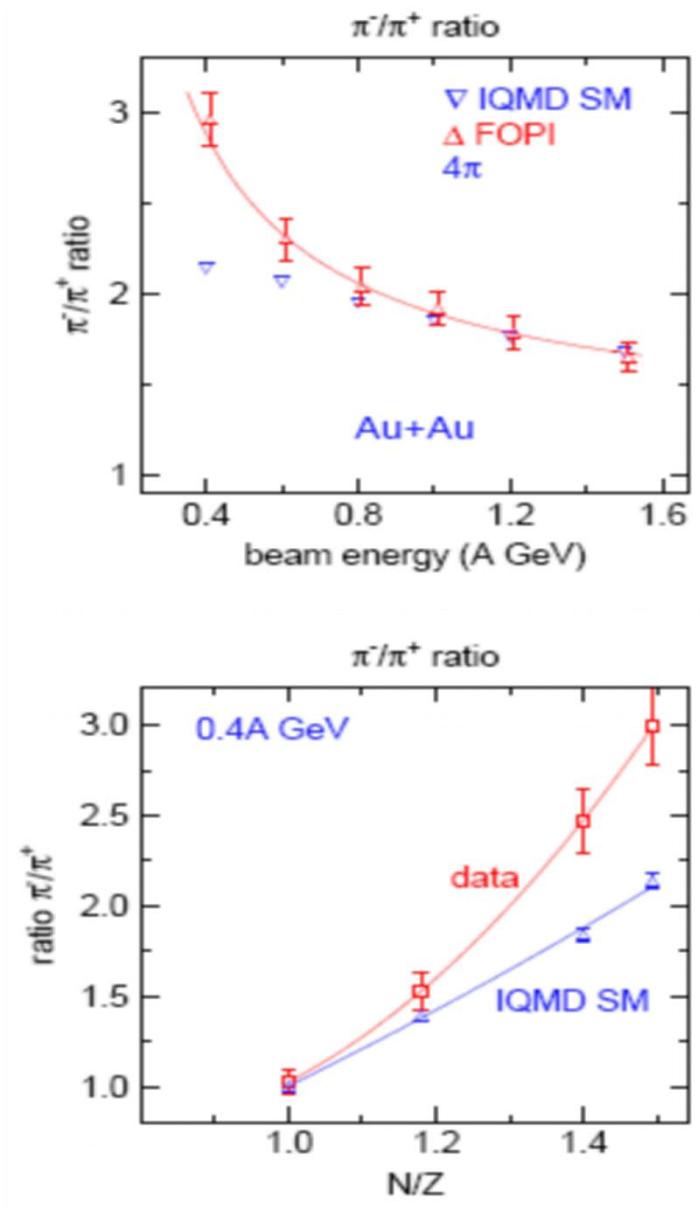
C. Fuchs et al., **PRL** 86, 1974 (2001)

Insensitive in light system C+C; Higher sensitivity found at lower energies



KaoS Collaboration

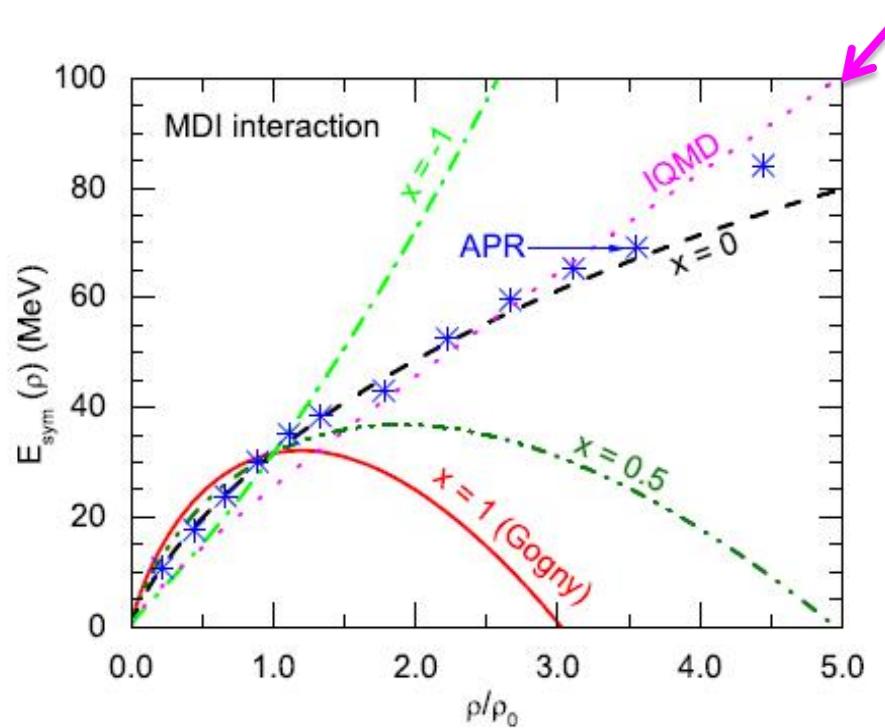
# $E_{\text{sym}}(\rho)$ : $\pi^-/\pi^+$ 产额比与 $E_{\text{sym}}(\rho)$



IQMD: C. Hartnack et al, EPJA 1 (1998) 151

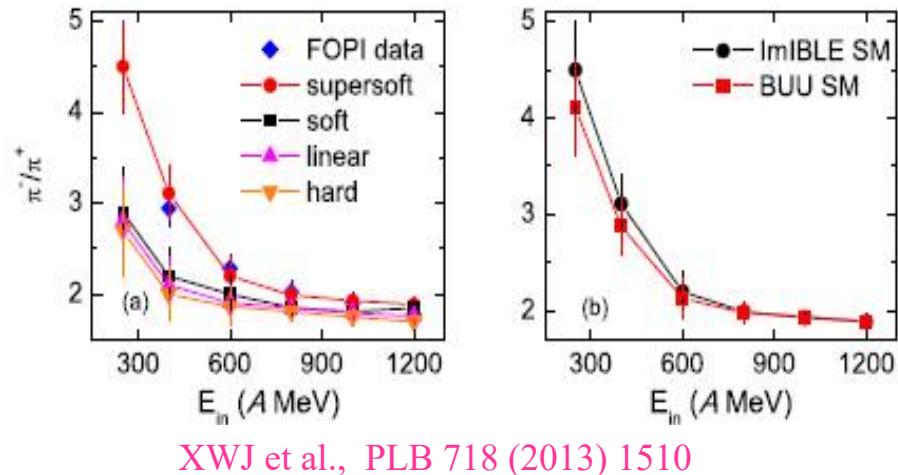
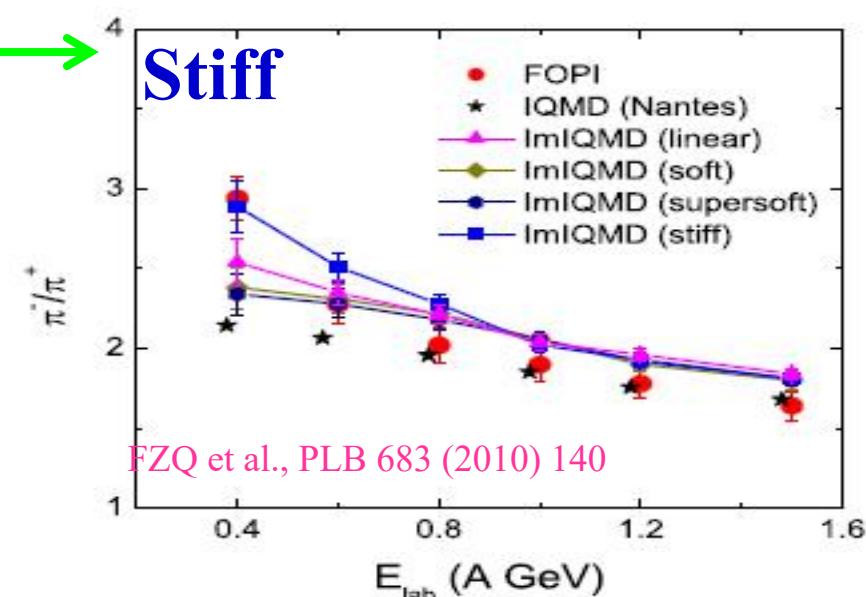
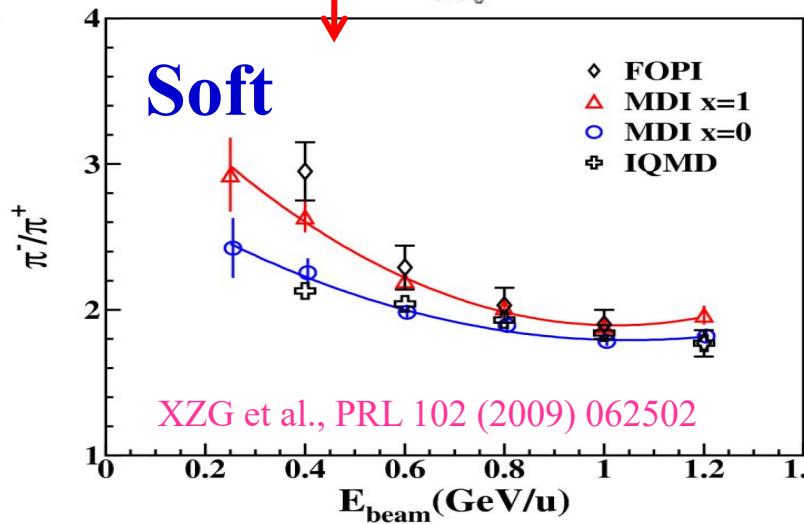
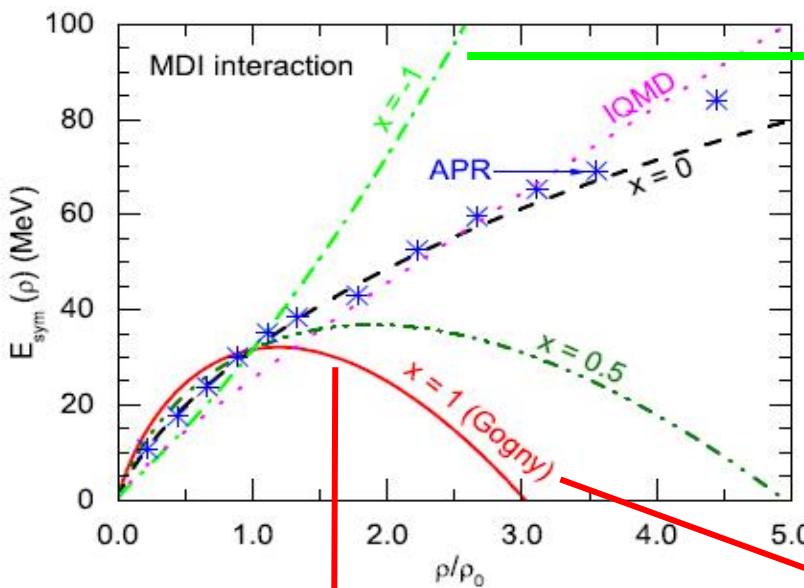
$$V_{\text{sym}}^{ij} = t_6 \frac{1}{\rho_0} T_{3i} T_{3j} \delta(\vec{r}_i - \vec{r}_j), \quad t_6 = 100 \text{ MeV}$$

$$\text{对应于对称能 } E_{\text{sym}}(\rho) = \frac{25}{2} \frac{\rho}{\rho_0} + (2^{2/3} - 1) \frac{3}{5} E_F^0 \left( \frac{\rho}{\rho_0} \right)^{2/3}$$



Need a softer symmetry energy to make the pion production region more neutron-rich!

# $\pi^-/\pi^+$ 产额比探针的模型依赖

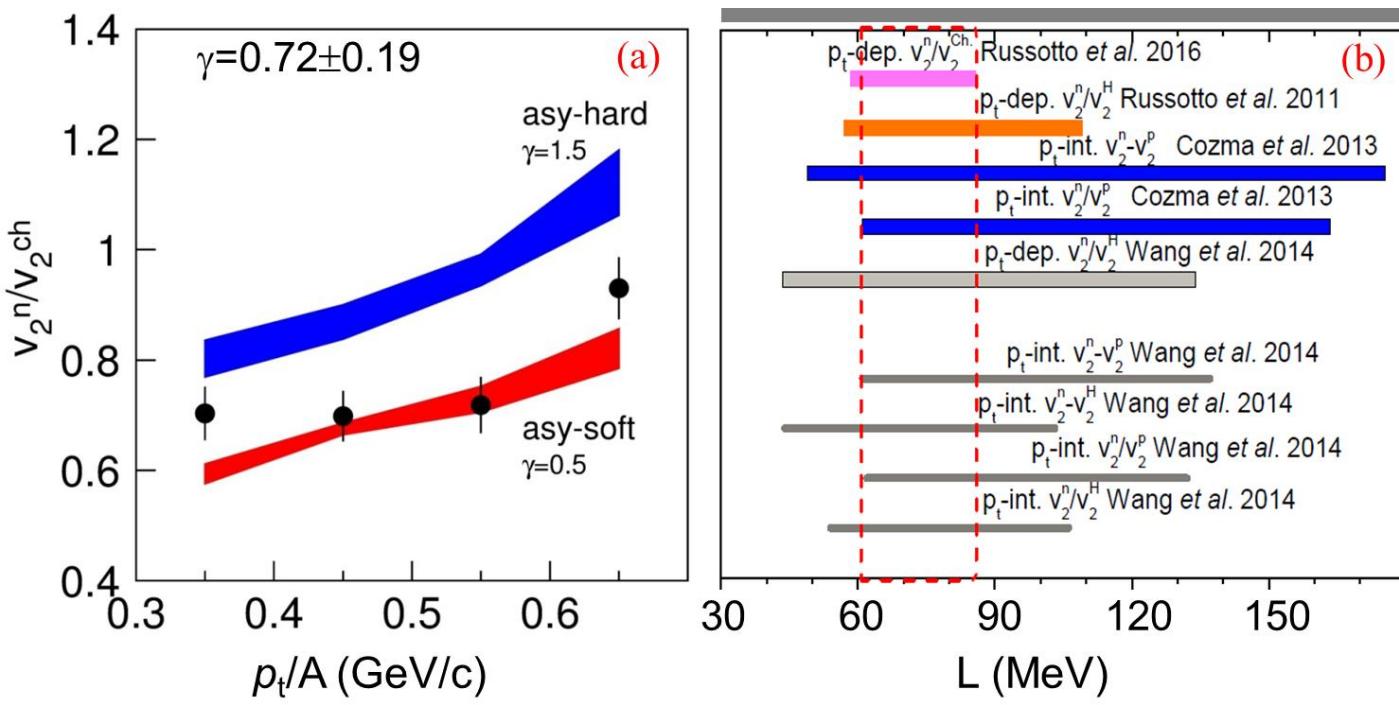
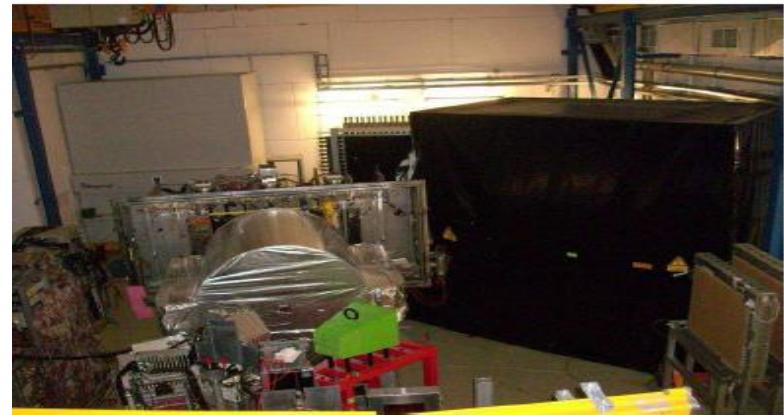
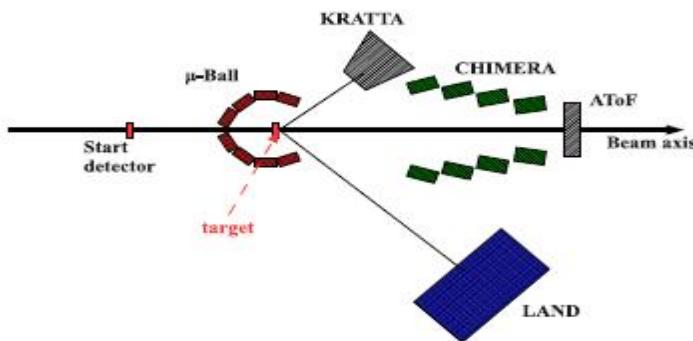


or loss sensitivity on  $\gamma$

J. Hong et al, PRC 90,024605(2010)

# 从 n/p 微分流提取 $E_{\text{sym}}(\rho)$

AnSYS-EOS 合作组



Y. Leifels et al  
PRL 71, 963 (1993)

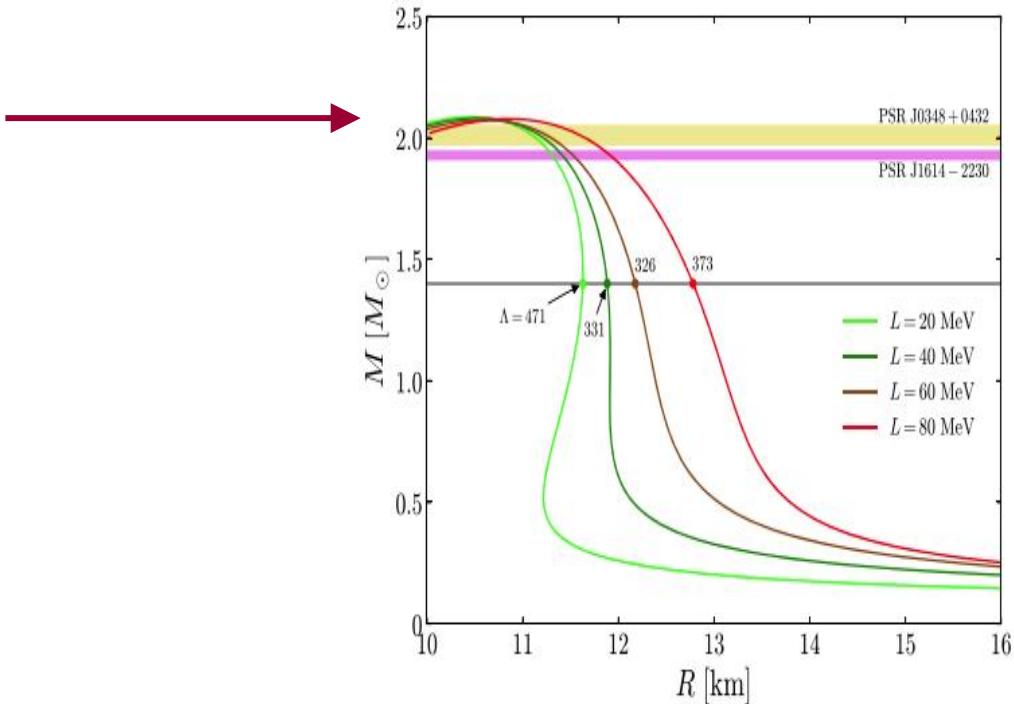
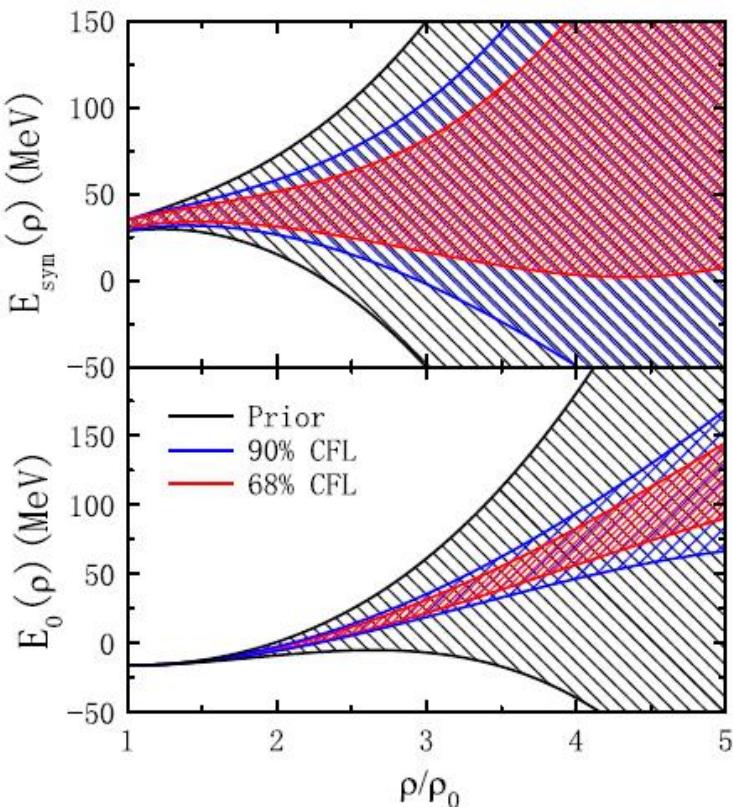
Russotto et al.,  
PLB 697, 471 (2011);  
PRC 94, 034608 (2016)

$$E_{\text{sym}}(\rho) = 12 (\rho/\rho_0)^{2/3} + 22 (\rho/\rho_0)^\gamma$$

# 从 GW170817 /PSRJ0740+6620 和 HIC 提取 $E_{\text{sym}}(\rho)$

$R_{1.4} = 11.9^{+1.4}_{-1.4} \text{ km}$ , PRL, 121, 161101, 2018;  $R_{1.4} = 10.8^{+2.1}_{-1.6} \text{ km}$ , PRL, 121, 091102, 2018

1. QMF18:  $L = 40 \text{ MeV}$
- Z. Y. Zhu et al., APJ 862, 98 (2018)



2. 贝叶斯分析,
- $E_{\text{sym}}(2\rho_0) = 39.1^{+12.1}_{-8.2} \text{ MeV (1}\sigma)$

W. J. Xie et al., APJ 883, 174 (2019)

N. B. Zhang et al., APJ 879, 99 (2019)

# 从 GW170817 /PSRJ0740+6620 和 HIC 提取 $E_{\text{sym}}(\rho)$

## 3. 联合中子星与核物理的数据

Y. Zhou et al., PRD99 121301(R) (2019)



$$L(N_c) = 47.3 \pm 7.8 \text{ MeV}$$

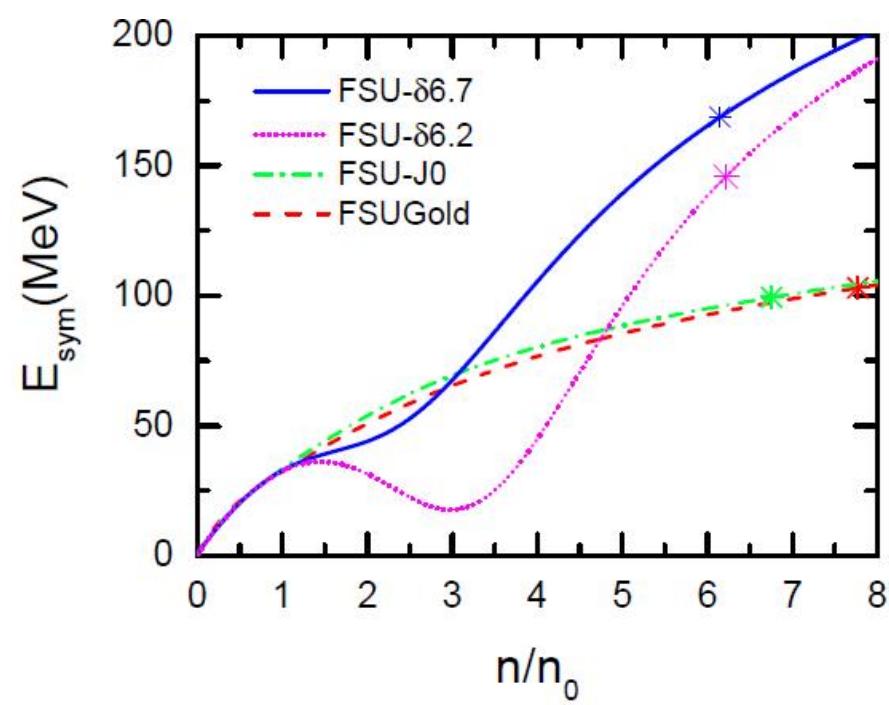
$$E_{\text{sym}}(2\rho_0) = [39.4^{-6.4}_{+7.5}, 54.5^{-3.2}_{+3.1}] \text{ MeV}$$

## 4. 考虑 $\delta - \sigma$ 耦合，提出两组新的RMF参数 **FSU-86.7, FSU-86.2**

F. Li et al., ApJ 929, 183 (2022)

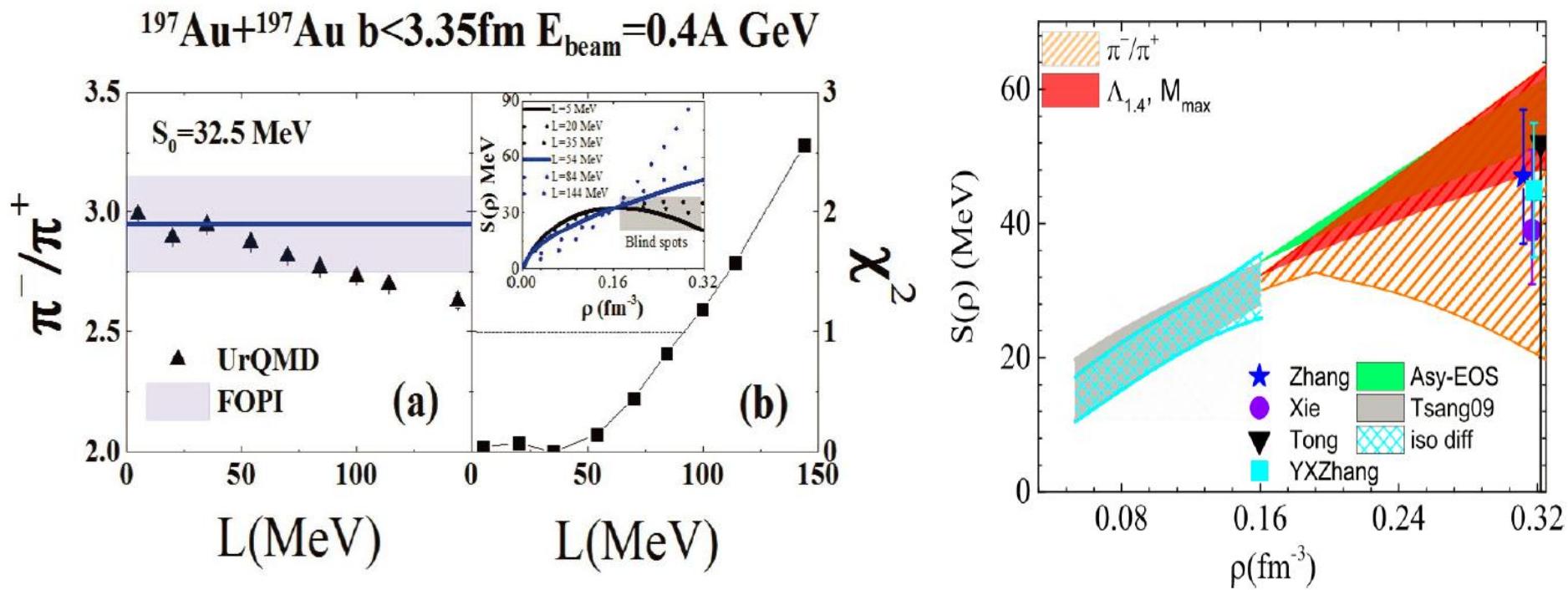
饱和点斜率分别为：

$$L(\rho_0) = 53.5 \text{ MeV}, 48.2 \text{ MeV}$$

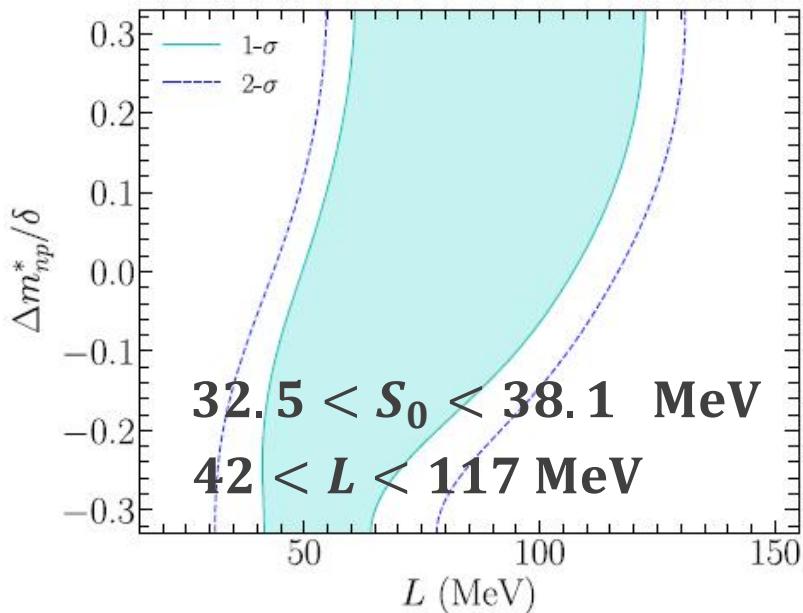
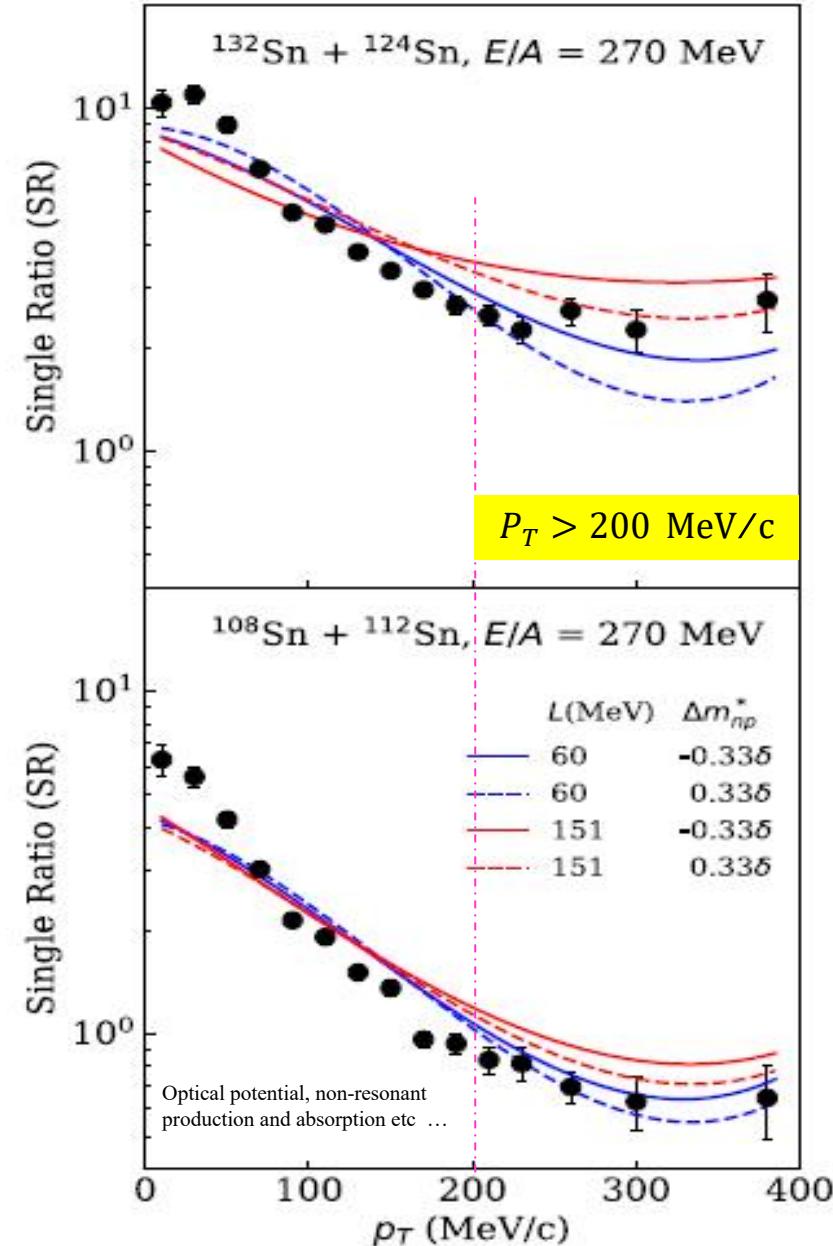


# 重离子碰撞和天体观测的结合

$\pi^-/\pi^+$  的复杂性，引发了大量对重离子反应中  $\pi$  介子产生和输运的研究，包括阈能效应、介质效应， $\pi$  介子光学势， $\pi$ - $\Delta$  循环等：徐骏，张英逊，C-M Ko, Dan Cozma 等等。



# S $\pi$ RIT 合作组最近进展

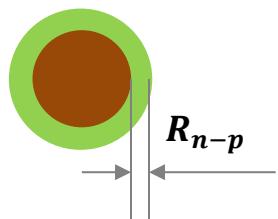


J. Estee et al., PRL 126, 162701 (2021)

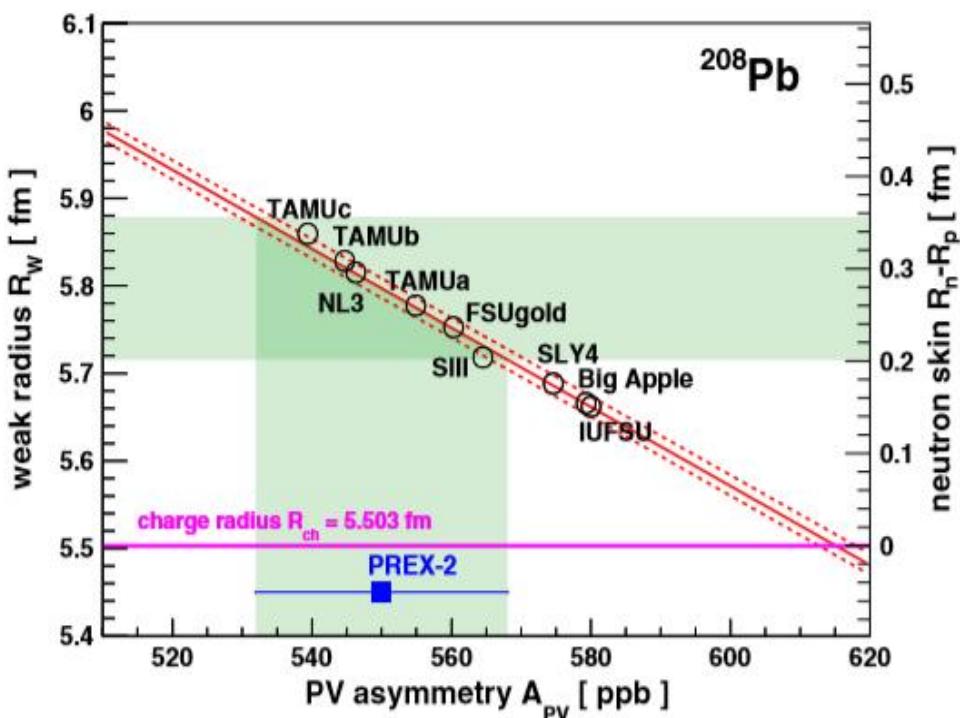


# PREX 实验的最新进展

$^{208}\text{Pb}$

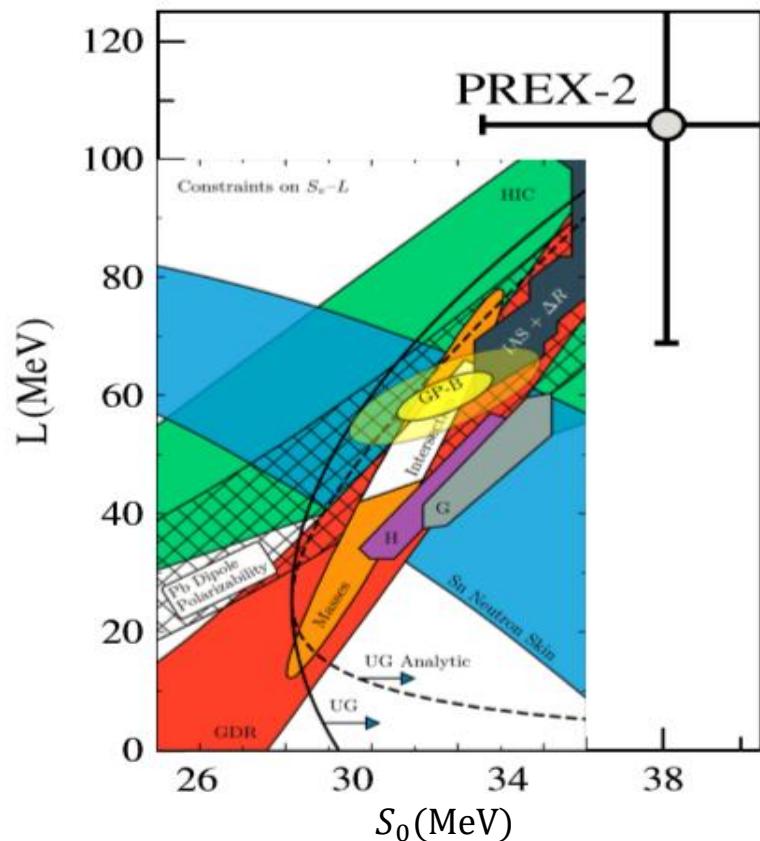


PREX 实验利用宇称破坏的  $e + ^{208}\text{Pb}$  散射，测量  $^{208}\text{Pb}$  中子皮。



$$R_{n-p} = 0.283 \pm 0.071 \text{ fm}$$

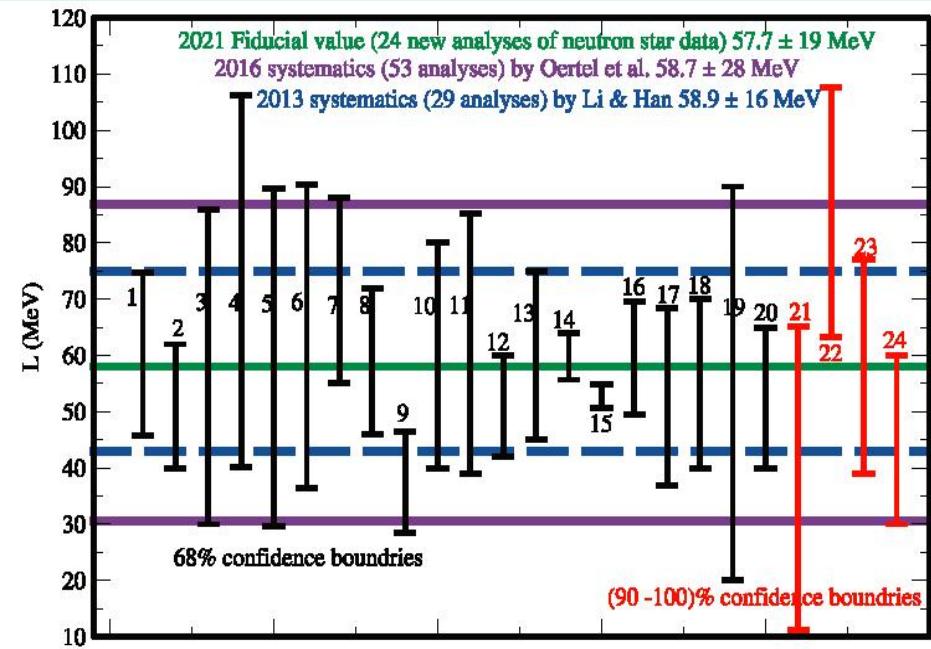
$$L = 106 \pm 37 \text{ MeV}$$



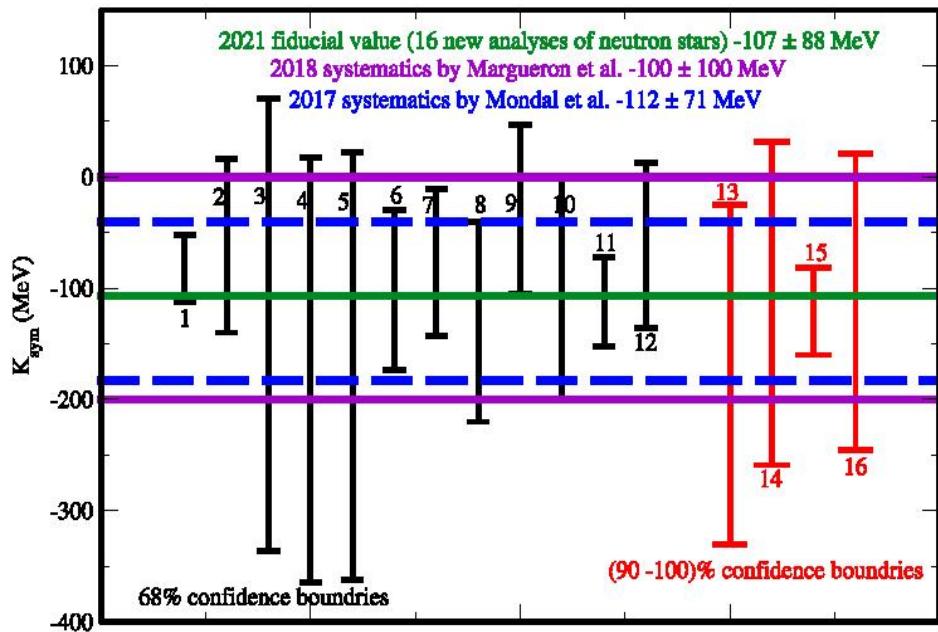
PREX collaboration., PRL 126, 172502 (2021)  
 B. T. Reed et al, PRL 126, 172503 (2021)

# $E_{\text{sym}}(\rho)$ 的置信结果

## Recent results on $L$



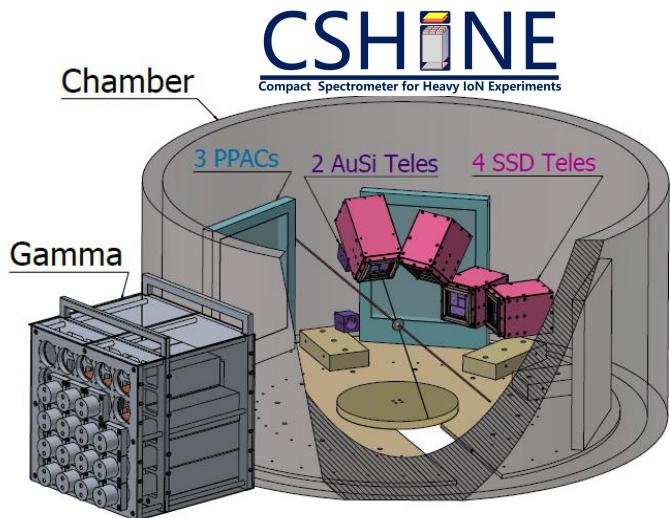
## Recent results on $K_{\text{sym}}$



Bao-An Li et al., Universe 7, 182 (2021)

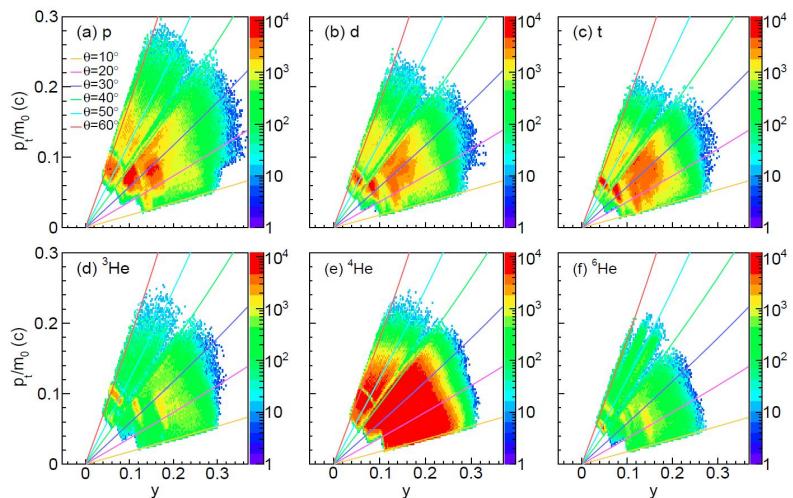
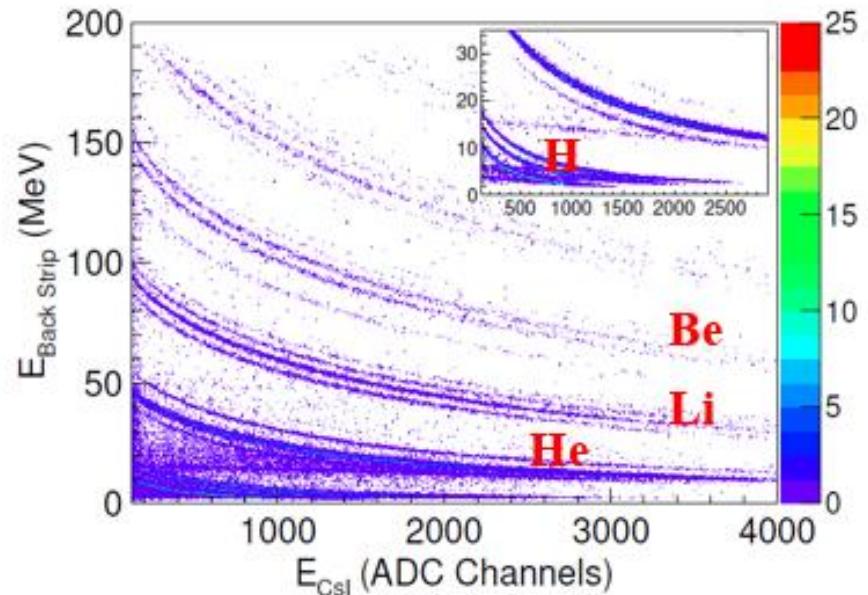
24 new analysis from the NS observables  
In comparison to 2013 and 2016 survey.  
In tension with the most recent PREX II results.

# CSHINE: Compact Spectrometer for Heavy Ion Experiment



NST 32, 4 (2021)  
NST 33, 40 (2022)  
NST 33, 162 (2022)  
NIMA 1011, 165592(2021)  
NIMA 1029, 166461(2022)  
NIMA 1053, 168330(2023)

PRC 95, 041602(R) (2017)  
PLB 825, 125856 (2022)  
PRC107, L041601 (2023)



Future plan: n/p, t/3He, Direct  $\gamma$  in HIC...

# CSHINE 实验在饱和密度附近的 $E_{\text{sym}}(\rho)$ 约束结果

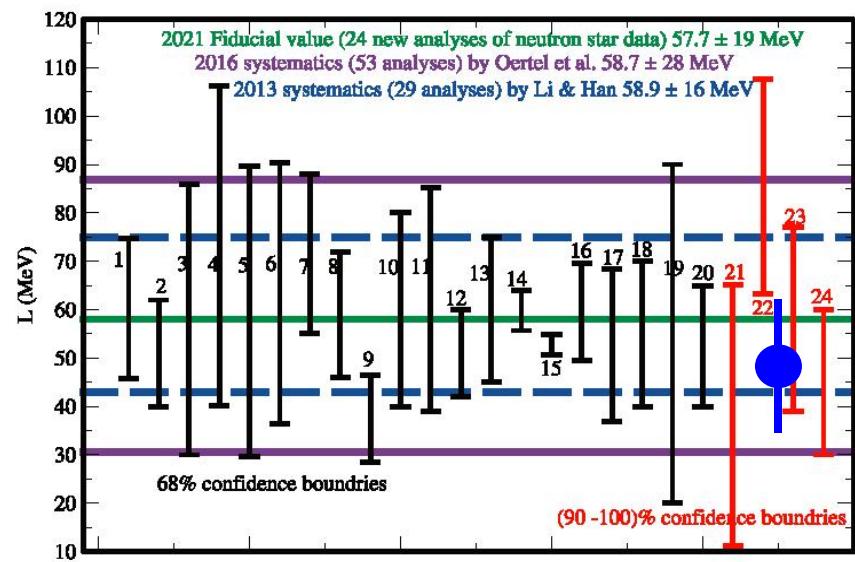
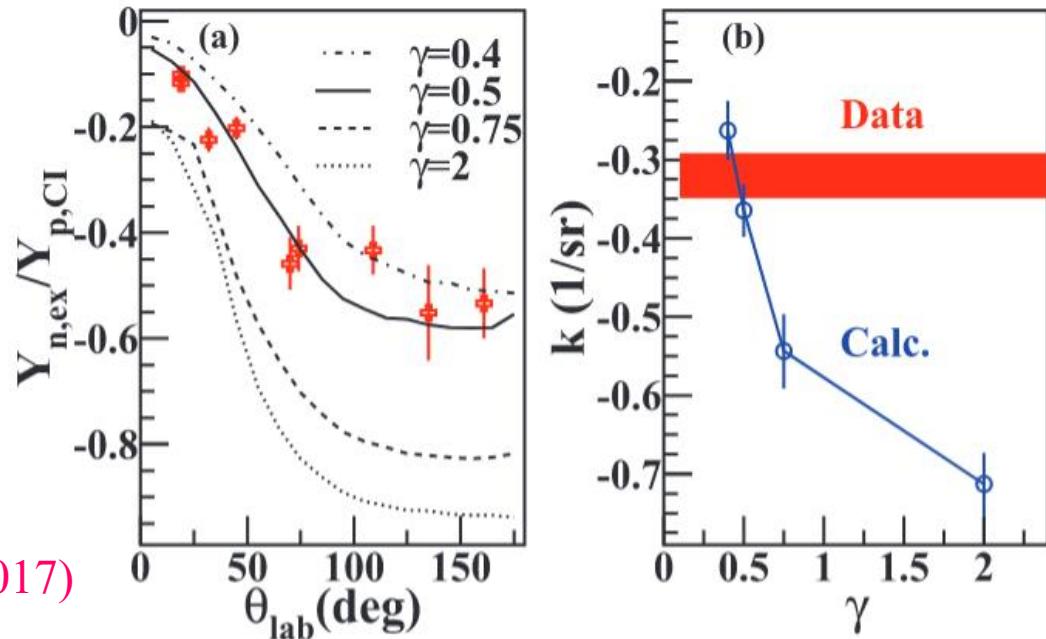
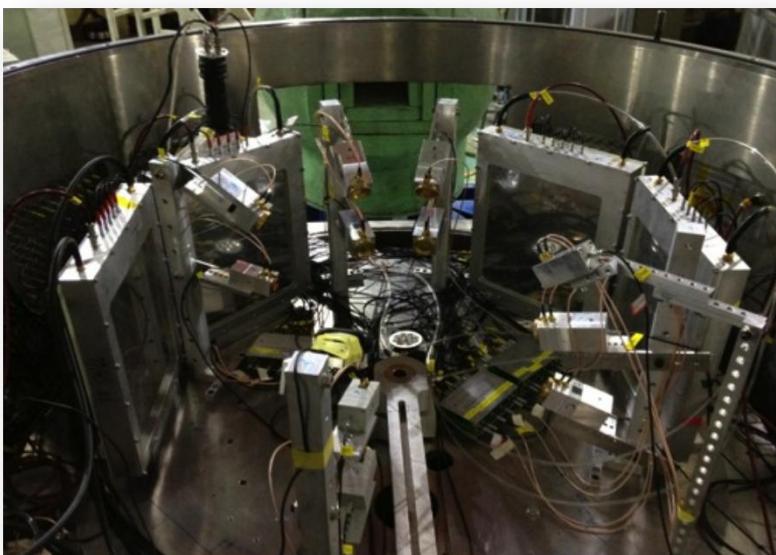
30 MeV/u  $^{40}\text{Ar} + ^{197}\text{Au}$

$E_{\text{sym}}(\rho)$ 约束结果:

$L = 47 \pm 14 \text{ MeV (95\% CL)}$

( $S_0 = 28.3 \text{ MeV}$ )

Y. Zhang , ... ZGX , PRC 95,041602(R) (2017)



## 二、CEE (HIRFL-CSR) 和 CEE+(HIAF) 的新机遇

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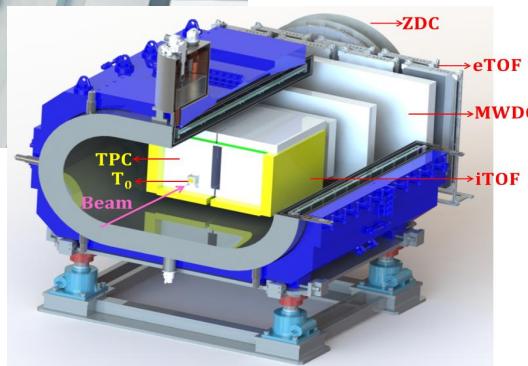
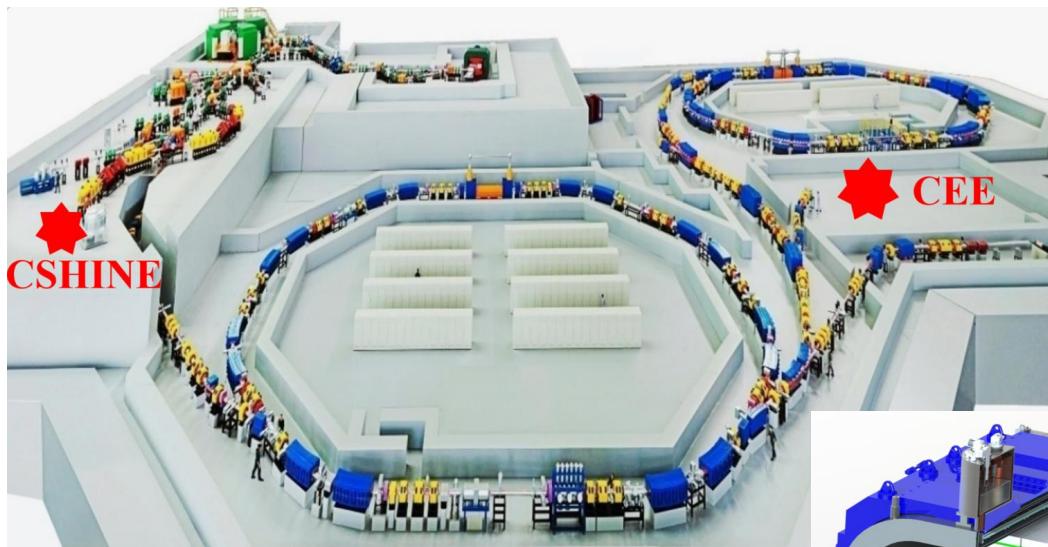
在饱和密度附近或者中等密度区，从 $\rho_0$ 到几倍 $\rho_0$ ，对应束流能量从数百MeV/u至数个GeV/u，已经有不少研究结果；

我们的研究动机：

是否还有其他机遇？

# 兰州低温高密核物质测量谱仪(CEE)

## HIRFL-CSR 加速器群



Cool Storage Ring



HIRFL-CSR:

P 2.8 GeV  
C 1100 MeV/u  
U 0.5 GeV/u

如果配备一个先进谱仪

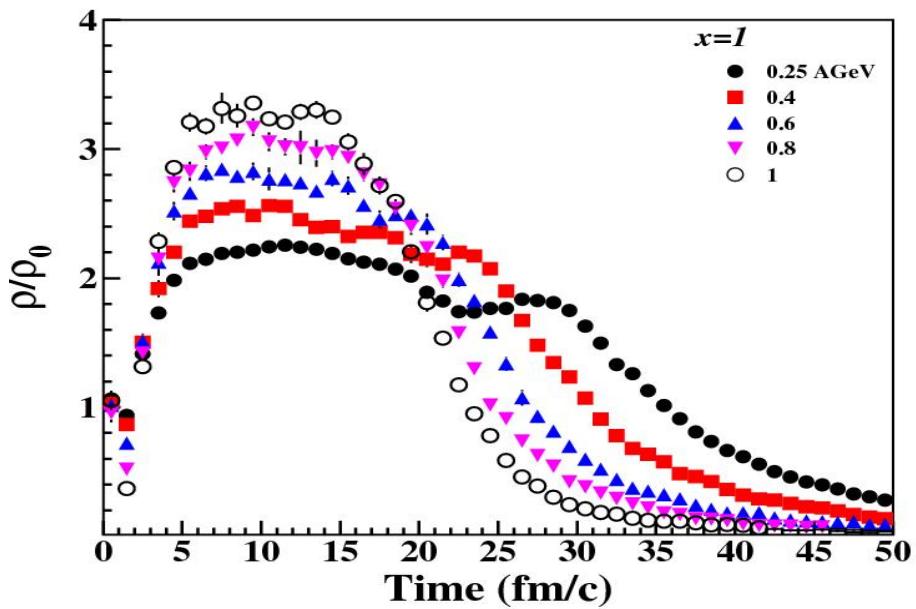
→ 寻找QCD相边界/高重子密度区EOS/相变临界点

→  $2\rho_0$  密度附近的EOS

→ 超核产生与作用势

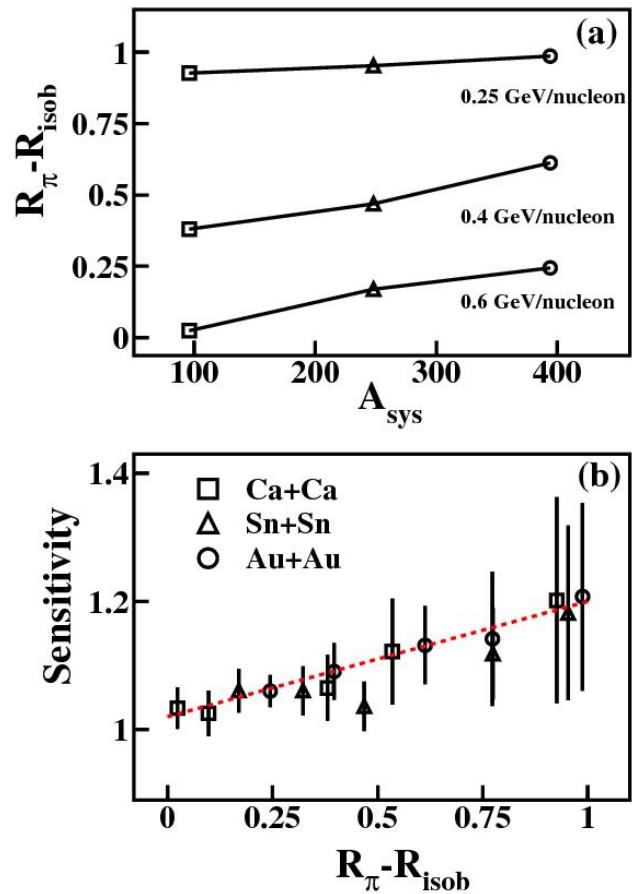
# HIRFL-CSR energy region is ideal for $E_{\text{sym}}(\rho)$ studies

- ↓ Nuclear matter at  $2\rho_0$  can be produced.  
→ The sensitivity of the observable on  $E_{\text{sym}}$  is enhanced.  
(Finite size effect, nuclear transparency)

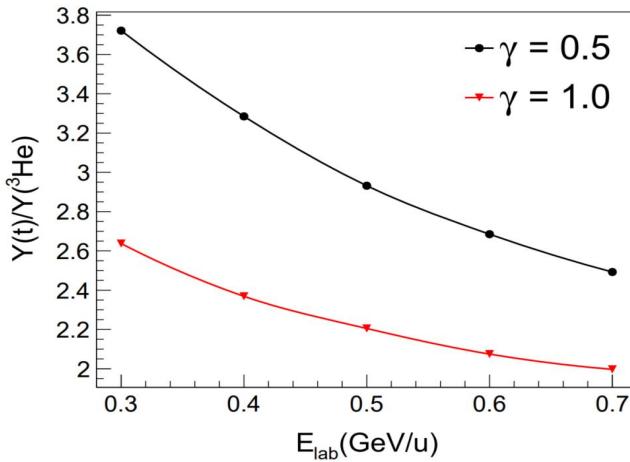
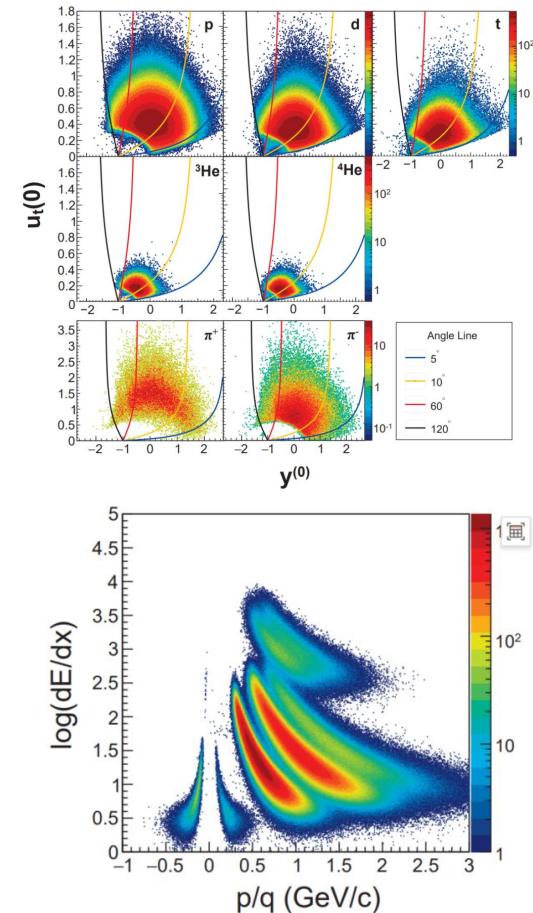
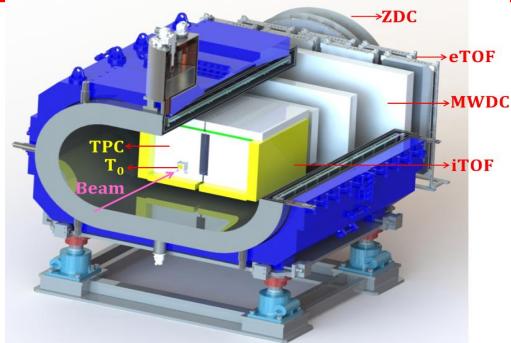


M. Zhang, ZGX et al., PRC 80 (2009) 034616

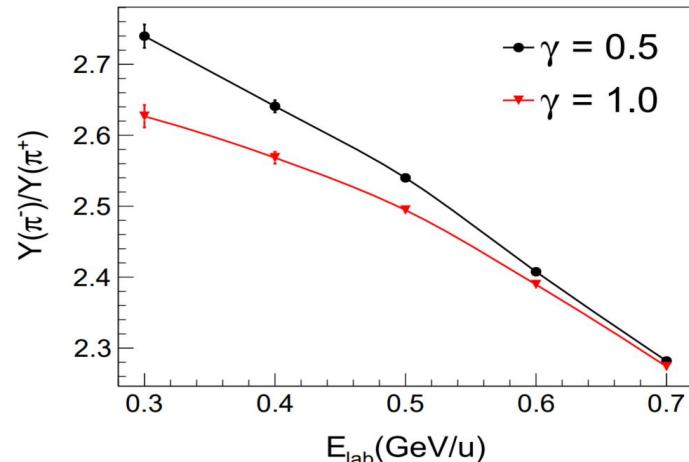
F. Fu ZGX et al, PLB 666 (2008) 359



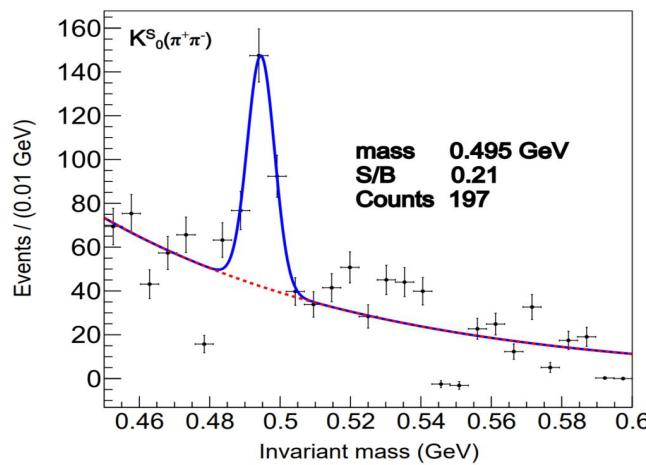
# CEE provides new opportunities for $E_{\text{sym}}(\rho)$ studies



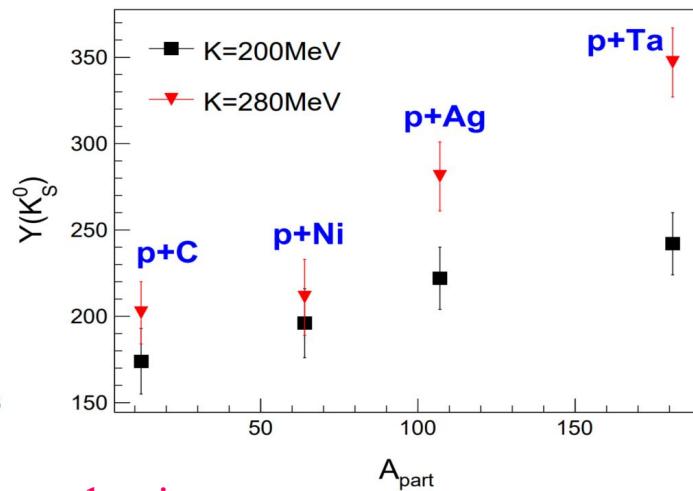
$t/{}^3\text{He}$  Ratio



$\pi^-/\pi^+$  Ratio



$K_s^0$  production

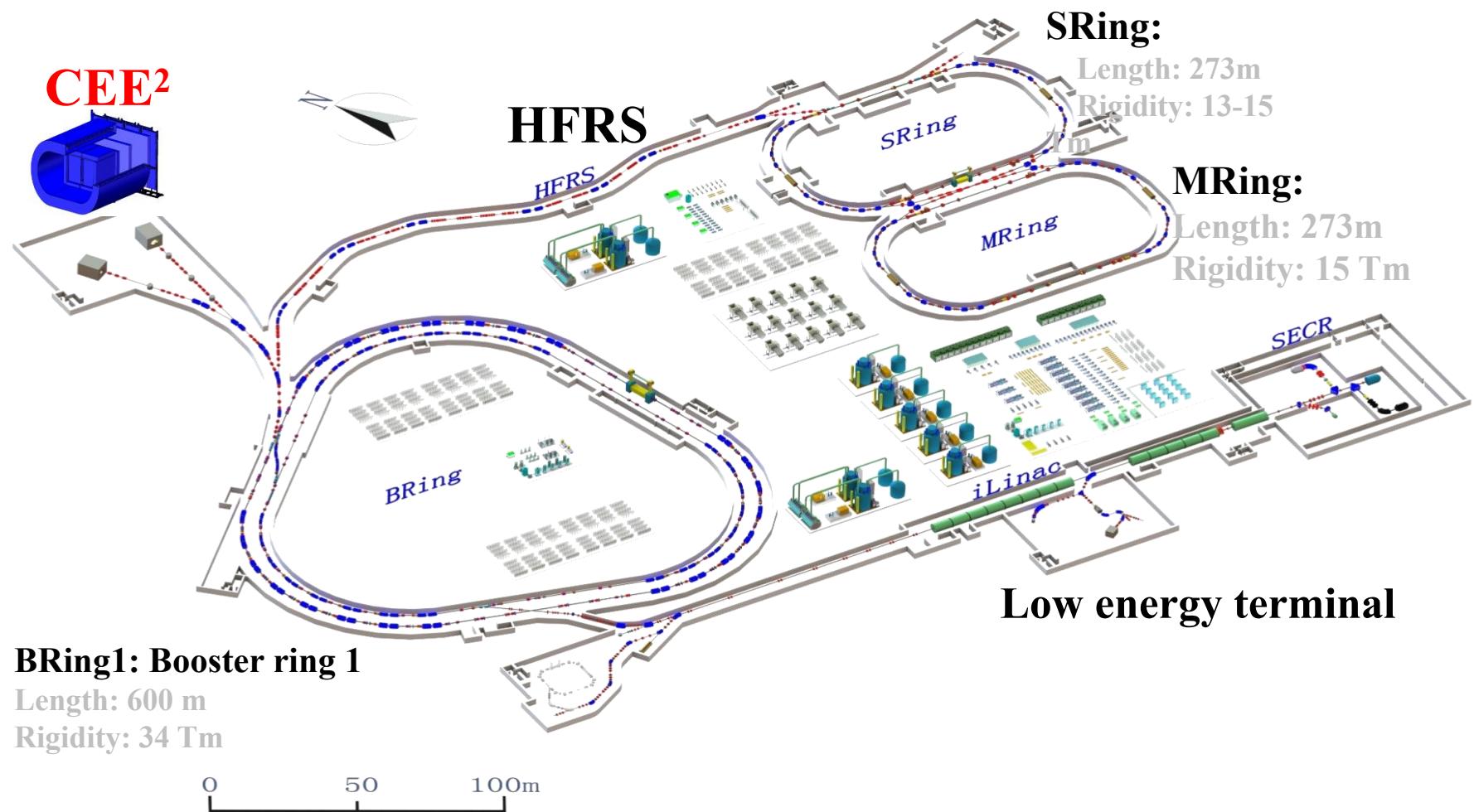


有关CEE的设计和研制进展，请听张亚鹏的报告

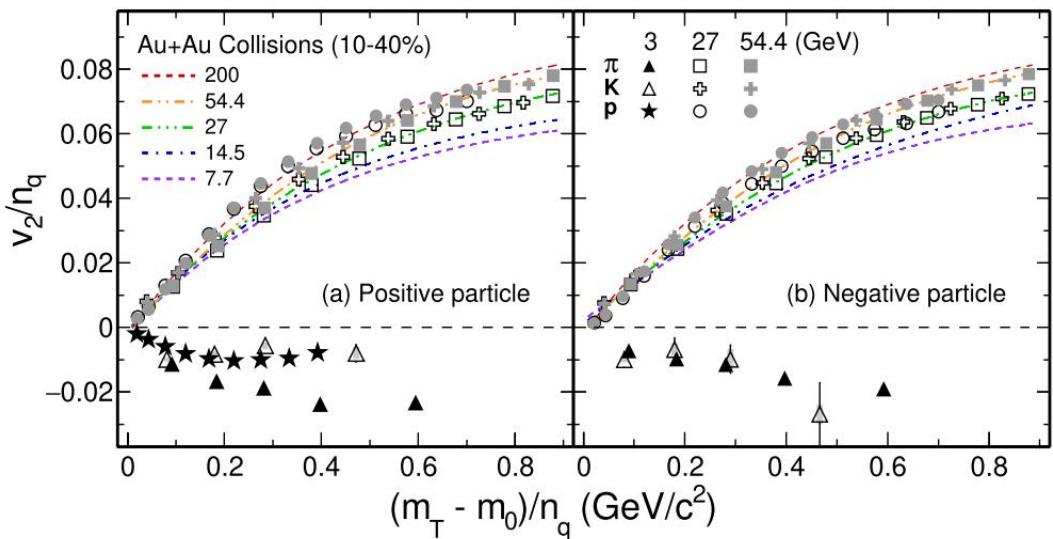
# 将来计划和机遇: HIAF

重离子束流能量(A/Z=2)  $\sim 4.5 \text{ GeV/u}$  <↔> 束流强度 HIAF/HIRFL:  $10^3 - 10^4$

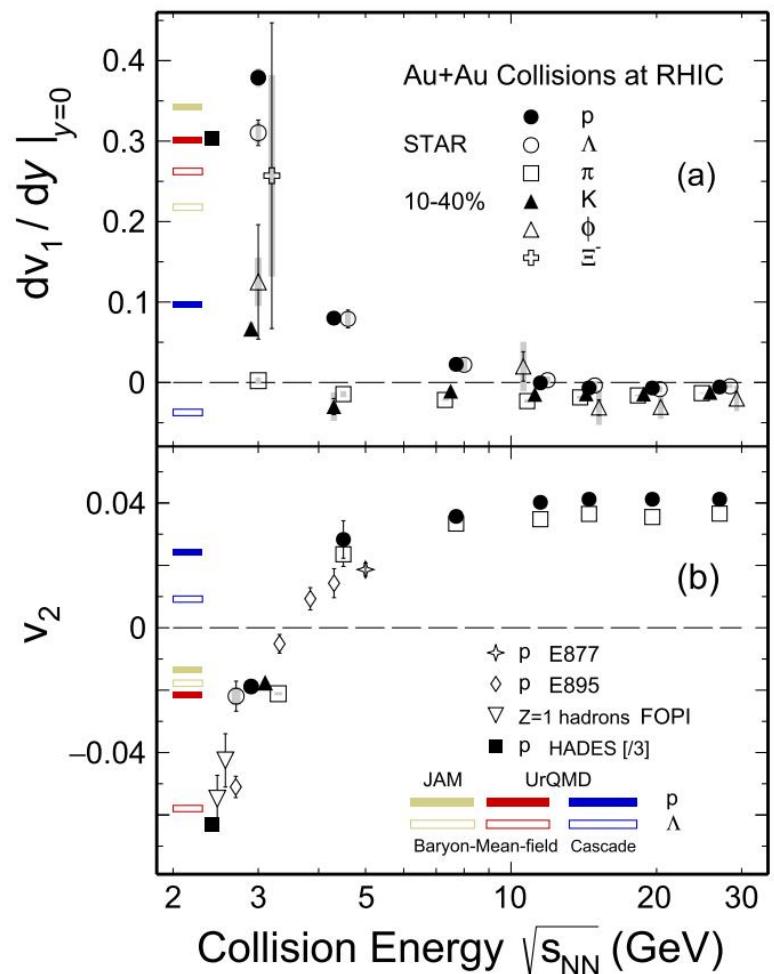
$$\rho \gtrsim 3 - 4\rho_0$$



# Recent results of STAR experiment $\sqrt{S_{NN}} = 3$ GeV



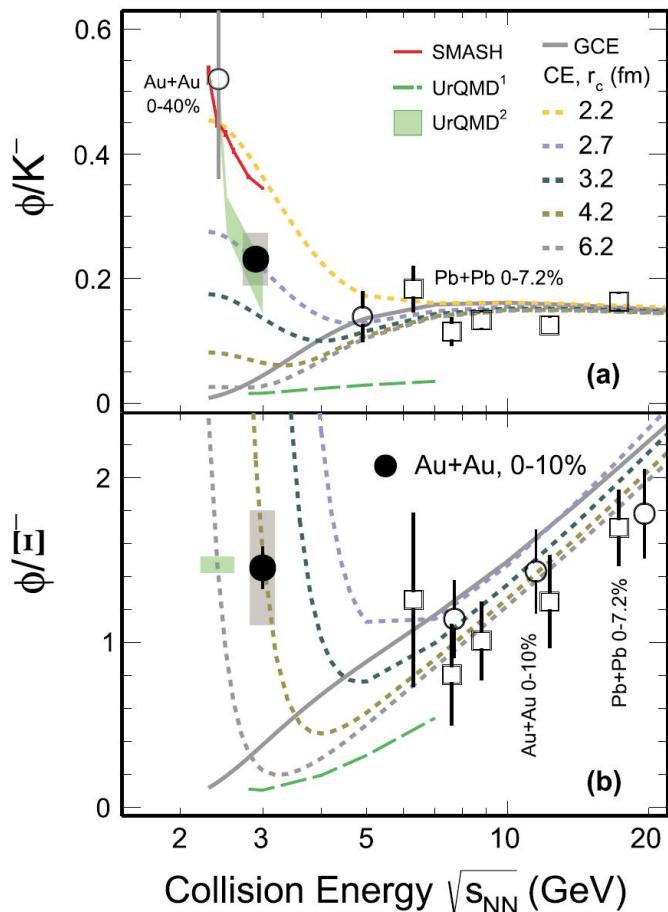
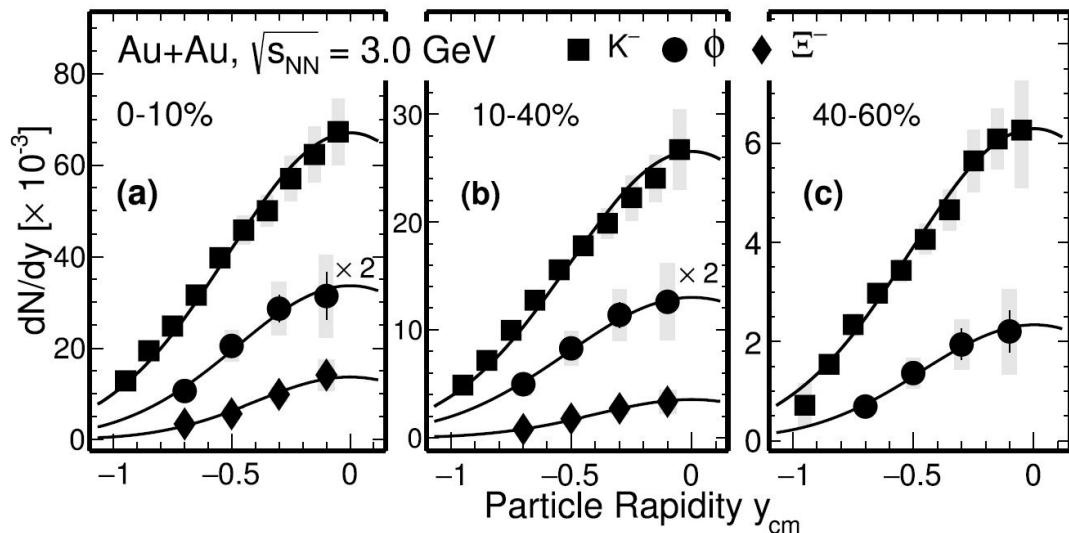
STAR Collaboration, PLB827, 137503 (2022)



At higher energy midcentral collisions the NCQ scaling holds, while at 3 GeV the  $v_2$  at midrapidity is negative and the NCQ scaling is absent, indicating the dense medium formed at  $\sqrt{S_{NN}} = 3$  GeV is likely hadronic matter.

# Recent results of STAR experiment $\sqrt{S_{NN}} = 3 \text{ GeV}$

STAR Collaboration, PLB827, 137003 (2022)



The production yields of  $K^-$ ,  $\phi$  and  $E^-$  demonstrate the necessity of the Canonical Ensemble for strangeness at low energy heavy-ion collisions. Hadronic transport model calculations (UrQMD and SMASH) including resonance contributions reproduce the data, suggesting a change of the medium properties at 3 GeV compared to those from higher energy collisions.

# 对称核物质EOS的新探针

当入射能量更高，例如在HIAF上，形成的核物质密度更高，有必要寻找更好的观测量。

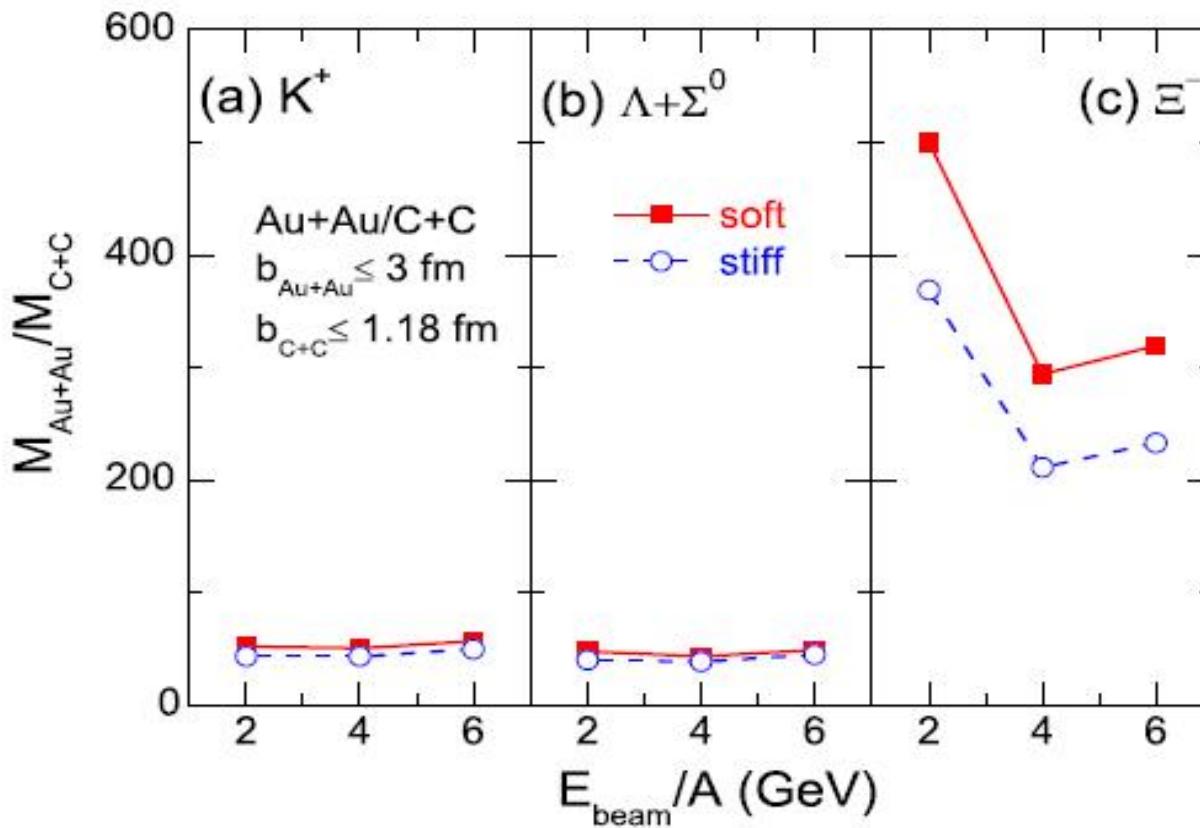
模拟发现， $\Xi^-$  通过次级反应、在更高的密度区域中才能产生，因此对高密区的核物质状态方程敏感。

为了研究中心密度  $\rho > 3\rho_0$  的状态方程，考虑2 GeV/u 以上的重离子碰撞，并考虑如下反应道；



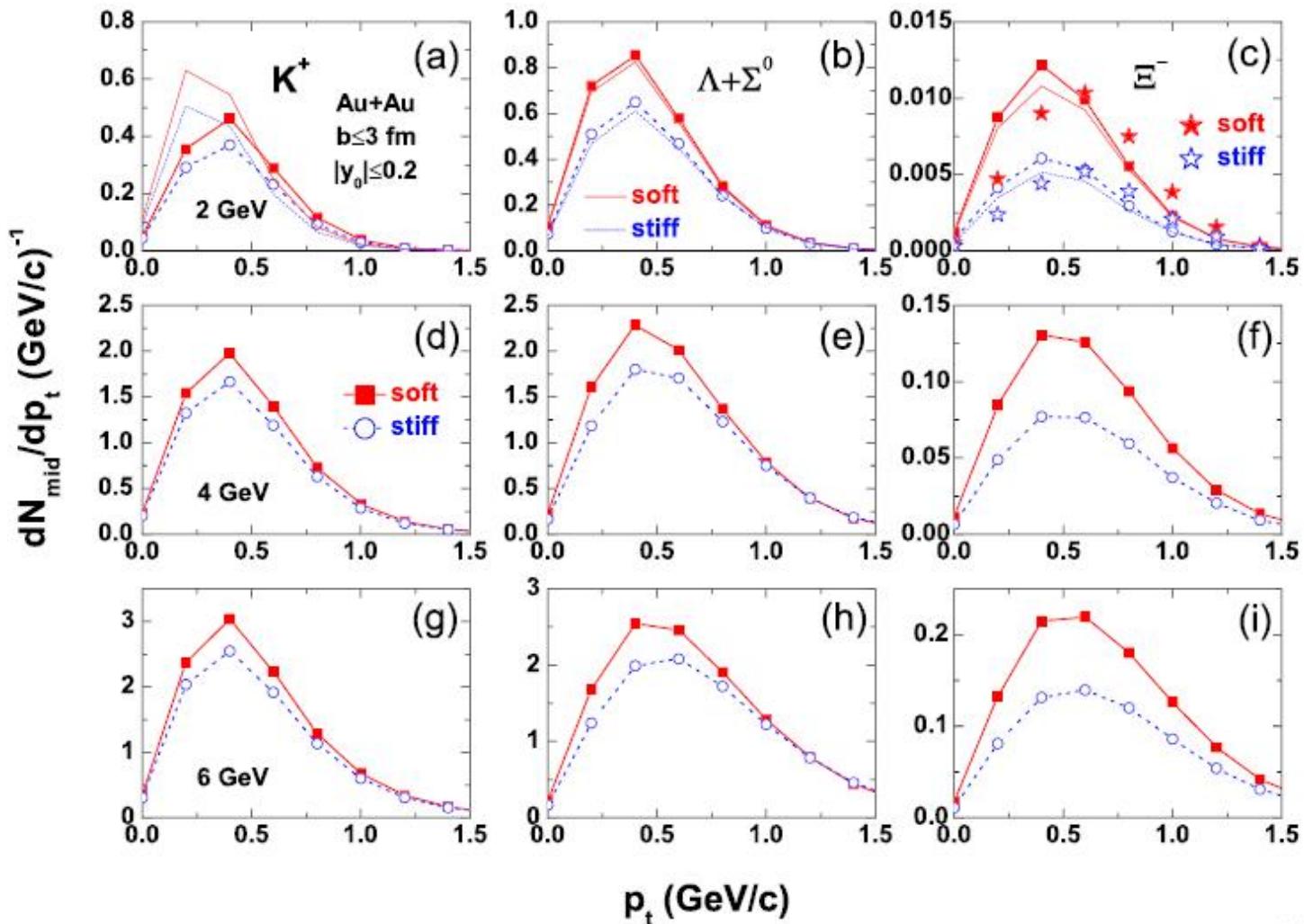
比较了两种不可压缩系数EOS的结果：  $K_0 = 200 \text{ MeV}$  和  $K_0 = 300 \text{ MeV}$

# 对称核物质EOS的新探针



从右图比较可以看出，随着入射能量升高，single strangeness 的敏感度减小了【与G. Ferini et al., PRL97, 202301 (2006) 的结论一致】，但是 double strangeness  $\Xi^-$  的产额，对EOS的依赖非常敏感。

# 对称核物质EOS的新探针



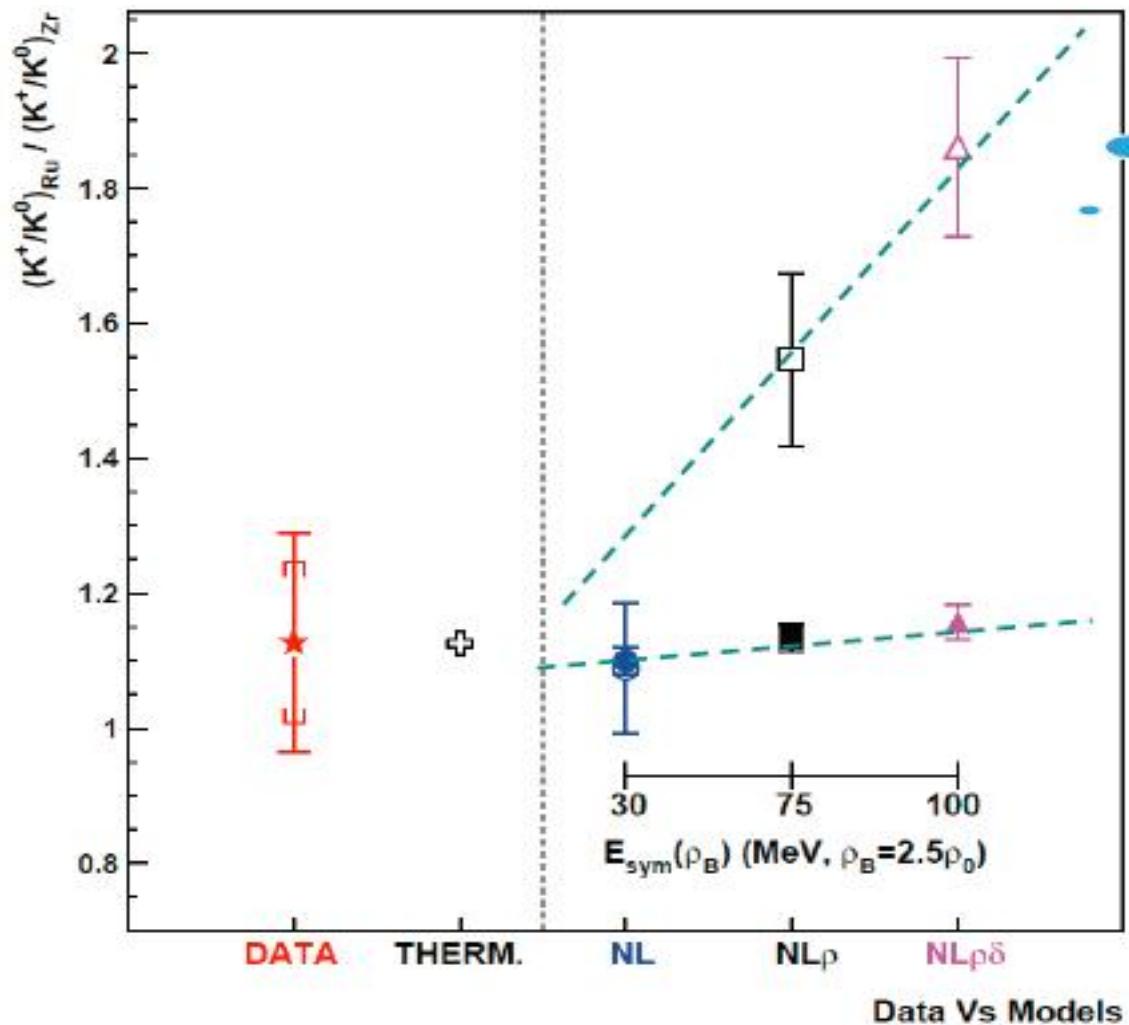
对EOS的依赖敏感度最大的范围，在 $p_t$ 分布的峰值处。

平均场和Strange hadrons的耦合，对 $\Lambda + \Sigma^0$  和  $\Xi^-$ 结果影响不大。

质量改变不是 $\Xi^-$ 产额对EOS的敏感依赖的原因。

# 更高密度区 $E_{\text{sym}}(\rho)$ 的探针

同位旋探针:  $\frac{n}{p} \rightarrow \frac{\pi^-}{\pi^+} \rightarrow \frac{K^0(d\bar{s})}{K^+(u\bar{s})}$



→ Collision dynamic effect reduces the sensitivity

→ Experimental precision not high enough to pin down  $E_{\text{sym}}(\rho)$

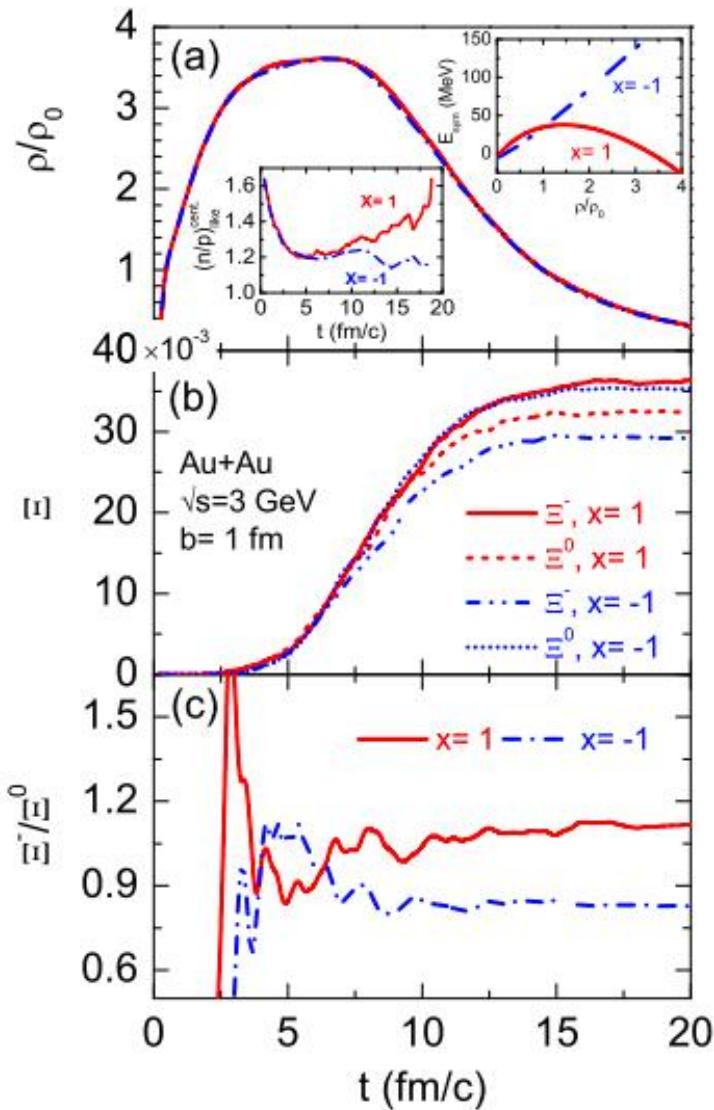
→  $NN \rightarrow N\Delta^-$  cancels out part of the N/Z asymmetry;

FOPI collaboration, PRC 75, 011901 (2007)

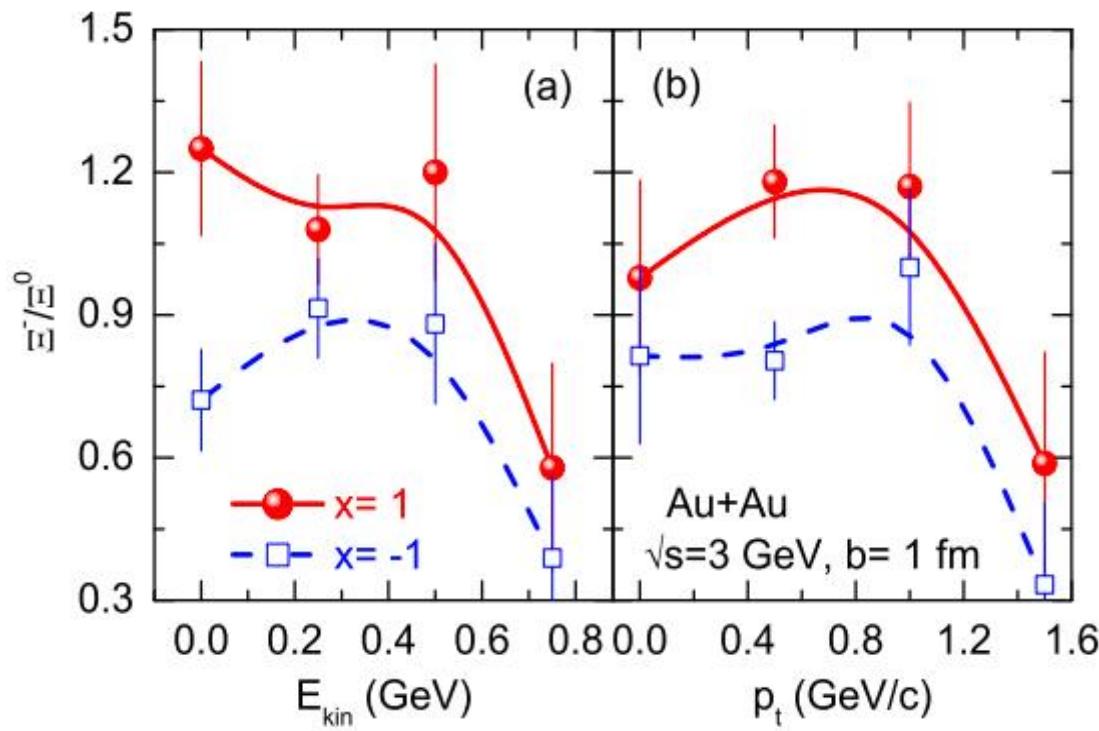
更重的系统，更大的N/Z跨度  
优化的束流能量

# 更高密度区 $E_{\text{sym}}(\rho)$ 的探针

同位旋探针:  $\frac{n}{p} \rightarrow \frac{\pi^-}{\pi^+} \rightarrow \frac{K^0(d\bar{s})}{K^+(u\bar{s})} \rightarrow \frac{\Sigma^-(dds)}{\Sigma^+(uus)} \rightarrow \frac{\Xi^-(dss)}{\Xi^0(uss)}$

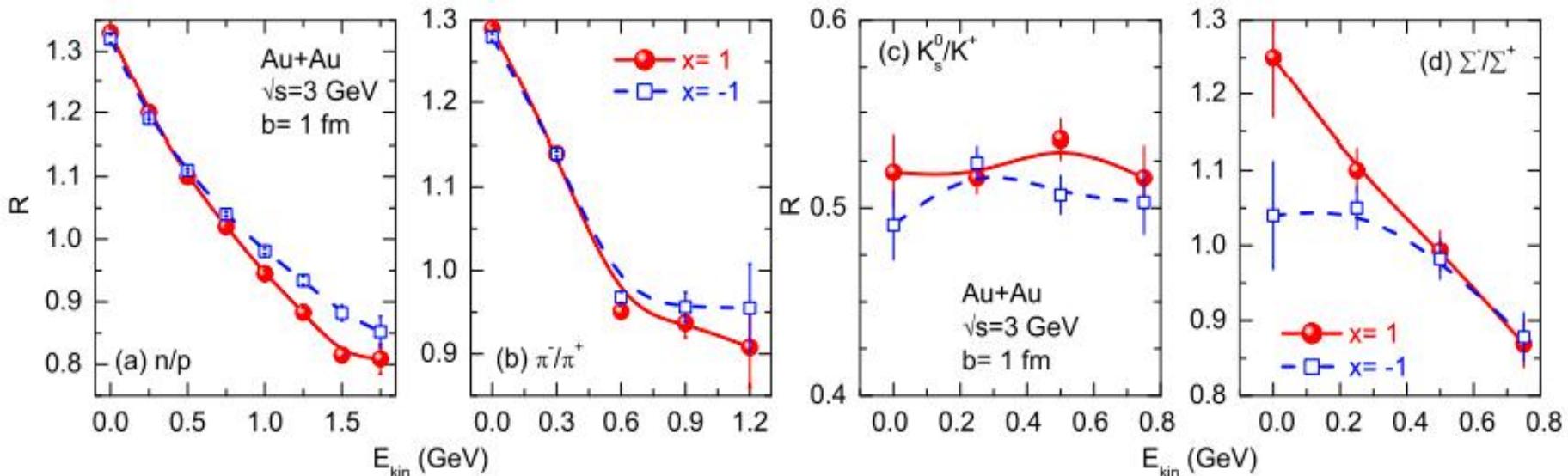


束流能量增加



# 更高密度区的对称能 $E_{\text{sym}}(\rho)$ 的约束

同位旋探针:  $\frac{n}{p} \rightarrow \frac{\pi^-}{\pi^+} \rightarrow \frac{K^0(d\bar{s})}{K^+(u\bar{s})} \rightarrow \frac{\Sigma^-(d\bar{s}s)}{\Sigma^+(uus)} \rightarrow \frac{\Xi^-(dss)}{\Xi^0(uss)}$



via  $\pi + N \rightarrow \Sigma$  secondly process, isospin effect carried by pions are passed to strangeness.

**实验难点:** 需要对末态中性粒子的重建和鉴别

$$\begin{aligned}\Sigma^- &\rightarrow \textcolor{magenta}{n} + \pi^-, \quad \Sigma^+ \rightarrow \textcolor{magenta}{n} + \pi^+ \text{ or } p + \textcolor{magenta}{\pi^0} \\ \Xi^- &\rightarrow \Lambda + \pi^-, \quad \Xi^0 \rightarrow \Lambda + \textcolor{magenta}{\pi^0}\end{aligned}$$

### 三、小结

重离子碰撞产生短暂的、微量的、压缩状态的“核物质”，因此是研究核物质状态方程(nEOS)的一个自然的手段！nEOS的研究，也是世界几个大型核物理实验室投放力量的方向。

基于国内大科学装置HIRFL，也可以产生一些成果（例如CSHINE上的相关工作）。

CEE 将在 GeV/u能区，为nEOS的研究，提供一些新的机遇。可测的物理量有轻碎片、 $\pi^-/\pi^+$  和  $K^0$ 介子等。

HIAF的建成，将在更高密度区( $\rho \gtrsim 3\rho_0$ )，为nEOS研究提供很好的研究机遇。在这一能区，多奇异子重子是很好的探针。

- 可以用重系统中的  $\Xi^-$  产额探测对称核物质不可压缩系数；
- 可以用  $\Xi^-/\Xi^0$ ,  $\Sigma^-/\Sigma^+$  等奇异探针约束对称能。

